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OFTICAL SIGNATURES PROGRAM FINAL REPORT

ROCKET PLUME OPTICAL SIGNATURES

October 1972

Prepared by:

McDonnell Douglas Astronautics Company-West 5301 Bolsa Avenue Huntington Beach, California 92647

for



CORPORATION

WASHINGTON OPERATIONS
1501 WILSON BLVD, SUITE 700, ARLINGTON, VA 22209



This report was prepared by McDonnell Douglas Astronautics Company-West under Subcontract No. 040-71-10 with the General Research Corporation, Washington Operations, for the U.S. Army Advanced Ballistic Missile Defense Agency under Prime Contract No. DAHC60-70-C-0078 MOD 5. Principal Investigator: M. Thomas (714) 896-4212. Optical Signatures Program Manager: T. M. Zakrzewski, (703) 524-7206, Extension 272.

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OPTICAL SIGNATURES PROGRAM FINAL REPORT ROCKET PLUME OPTICAL SIGNATURES

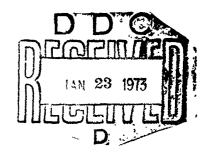
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PREFACE

This document is the final report on the "Rocket Plume Optical Signature Program." This work was performed for General Research Corporation, 1501 Wilson Boulevard, Suite 700, Arlington, Virgnia 22209, under Contract 040-71-10. This contract covered the period from 20 August 1971 to 30 September 1972.

The effort described herein was accomplished by the following study personnel: S. S. Cherry, M. Thomas, and R. L. Younkin. This report was approved by H. Hurwicz, Chief Advance Technology Engineer, Aero/
Thermodynamics and Nuclear Effects Research and Development, Advance Systems and Technology, McDonnell Douglas Astronautics Company-West.

ABSTRACT

As part of the Optical Signatures Program, McDonnell Douglas
Astronautics Company-West has developed the initial working model of a
code to describe the gross features of rocket-plume radiation for
altitudes above 75 n mi. The main effort is the construction of a
scheme for integration of an arbitrary function through an arbitrary
axisymmetric rocket plume, with any specified look angle, plume direction,
and vehicle velocity direction. Radiances are presented as integrated
values in a specified spectral band. The equations used and a printout
of the code and of a sample application are included.

CONTENTS

Section 1	INTRODUCTION AND SUMMARY			
Section 2	THEORY			
	2.2	Plume Gas Flow Particles Atmospheric Excitation	7 8 17	
Section 3	FLAME CODE			
	3. 1 3. 2	Geometry Considerations Flow Charts	24 33	
Section 4	EQUATIONS TO BE SOLVED			
	4. 1	Program Section 1. Initial Calculations — Particles	43	
	4.2	Calculations - Gas	45	
	4.3	Program Section 3. Selection of x ₁ or z ₁	46	
	4.4	Intercept Values	47	
	4.5 4.6	Radiance Calculations	47	
	4. 7	Program Section 6. Call SIMP— Integration Through Plume Curve Fitting Routine, CF	52 52	
Section 5		PLE PROBLEM AND OPERATING	J	
Section 5		RUCTIONS	54	
Section 6	RECOMMENDATIONS FOR FUTURE WORK			
Section 7	REFERENCES			
Appendix A	LISTING OF FLAME CODE			
Appendix B	OUTPUT FOR SAMPLE PROBLEM			
Appendix C	AUXILIARY LIBRARY ROUTINES			

FIGURES

1-1	Components and Flow of FLAME Code	3
1-2	Isointensity Lines for Broadside Look at Exatmospheric Plume of Solid Propellant Rocket	4
1-3	Isointensity Lines for Look at 45° from Plume Axis of Solid Propellant Rocket	5
1-4	Isointensity Lines for Broadside Look at Particle and Excited Water Radiation from Solid Propellant Rocket	6
2 - 1	Knudsen Number Variation with Engine Conditions	9
2-2	Plume Density Distribution on Spherical Cap (r = constant)	10
2-3	Particle Temperature, Flow Angle, and Velocity on Limiting Streamline	13
2-4	Alumina Particle Emissivity	18
2 - 5	Water Rotational Energy from Single Collision	19
2-6	Rotational Excitation Parameters for H ₂ O	20
2 - 7	Absorption Coefficient from Water Rotation at One Atmosphere Pressure	23
3 - 1	Plume Oricitation and Viewing Direction	25
3-2	Location of Vehicle Velocity Vector and Upward Direction in Coordinate System of Figure 3-1	34
3-3	Master Flow for Program P2170, FLAME	37
3 - 4	Program Region 4 Flow	38
3 - 5	Program Region 5 Flow	39
3-6	Program Region 6 Flow	40
3-7	Particle Emissivity Koutine Flow Chart	41
3-8	Absorption Coefficient Routine Flow Chart	41
3-9	Contour Plotter Flow Chart	42

TABLES

3-1	Listing of Input Data	35
3-2	Program Region 1. Initial Calculations- Particles	36
3 - 3	Program Region 2. Initial Calculations-Gas	36
3-4	Program Region 3. Select x_1 or z_1 for $\alpha \ge 90 - \theta_m$	36
4-1	Particle Limiting Velocity Coefficients	49
4_2	Concentration Coefficients	40

Section 1 INTRODUCTION AND SUMMARY

A code has been developed to describe the gross features of rocket-plume radiation for altitudes above 75 mmi (P2170-FLAME). The effort was intended to provide an initial working model, consistent with available technology. Theoretical studies were not intended to be conducted in any of the disciplines required. The intent was, however, to provide a suitable framework for the description of a plume with sufficient flexibility to incorporate the results of future research.

This objective was met by considering the two most important far-field radiative mechanisms, particle emission and atmospheric excitation, in their present state of understanding. With these major elements of plume radiation included even in crude form, future program modification to include revised theories should not require complete program revision. The main effort which has been pursued in FLAME-code development is the construction of a scheme for integration of an arbitrary function through an arbitrary axisymmetric rocket plume, with any specified look angle, plume direction, and vehicle velocity direction. The assumption of flow symmetry about the plume axis may prove poor for late-time plumes, but to include a complete three-dimensional capability would have burdened the code's development to preclude an operational version at the end of this contract. The switch to 3-D is not difficult at a later time and will involve obtaining flow correlations as a function of the azimuthal angle, as well as r and θ , and integrating over four quadrants instead of two.

Radiances are presented in the current code as integrated values in a specified spectral band. Spectral intensities are readily available for the particles, but only a band theory is available for the atmospheric excited species, due primarily to the species cooling characteristics. Again, a switch to spectral curves is possible with ease at a later time when a more general cooling theory evolves.

A schematic of the FLAME code is presented in Figure 1-1. Radiative cooling of both the hot particles in multiphase plumes and the excited molecular radiators is included. The four boxes feeding into the FLAME code represent the nain areas programmed during this contract. The use of a simplified flow model reduces both computer storage requirements and run time, and allows a more extensive checkout of the radiative mechanisms. More sophisticated flow models exist and could be substituted at a later time.

Input to the code includes simple engine parameters (Section 3, Table 3-1). Output is a plot of lines of constant radiance in the plume as viewed in the specified direction. A typical output from the code is shown in Figure 1-2 for a broadside look at a solid propellant plume in the absence of atmospheric excitation. The radiance is plotted in a plane perpendicular to the direction of view. The plume of Figure 1-2 is shown again in Figure 1-3 where the observer is at a 45-degree angle to the plume axis. Radiance from the same engine at low altitudes, including the effect of atmospheric excitation of the plume, is shown in Figure 1-4. The viewer angle, plume direction, and free stream velocity direction are all completely arbitrary. (Section 3, Figures 3-1 and 3-2.)

Only one-half the plume is treated by a single FLAME run. The remaining half can be obtained by redefining the azimuthal angles as shown in Section 3, Figure 3-2. It is not always necessary to generate both plume halves because the plume radiance is axisymmetric under certain conditions.

The radiative models used tend to underpredict the radiation. For atmospheric excitation, only single collisions are considered and the Boltzmann distribution of excited rotational states is assumed. Persistance of the plume radiance due to pumping by thermal (rather than pure kinetic) collisions with the atmosphere is not treated. For particle radiation, scattering of Earthshine is not treated.

Section 2 summarizes the theory used. Sections 3 and 4 detail the workings of the FLAME code, describing the equations used.

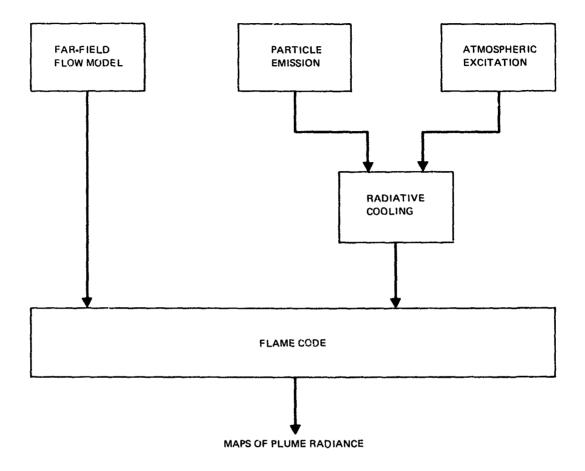


Figure 1-1. Components and Flow of FLAME Code

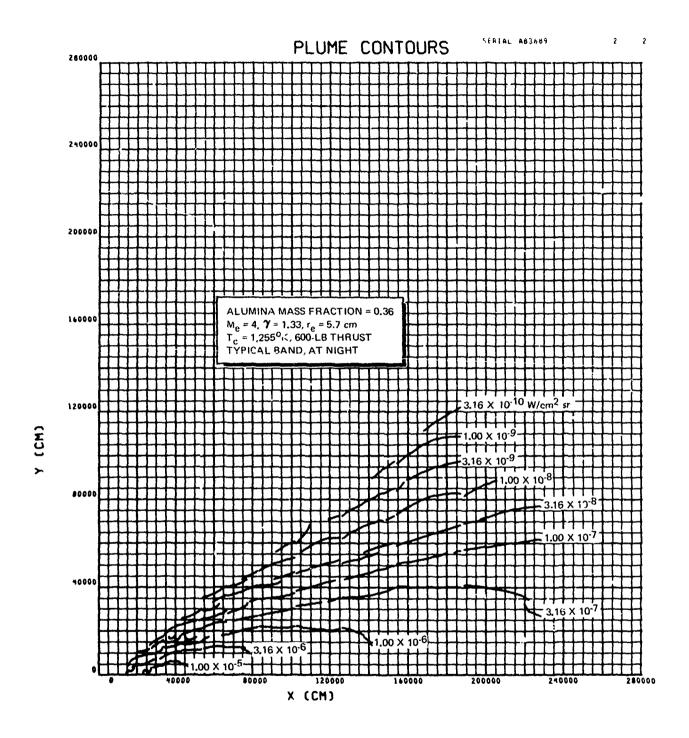


Figure 1-2. Isointensity Lines for Braodside Look at Exoatmospheric Plume of Solid Propellant Rocket

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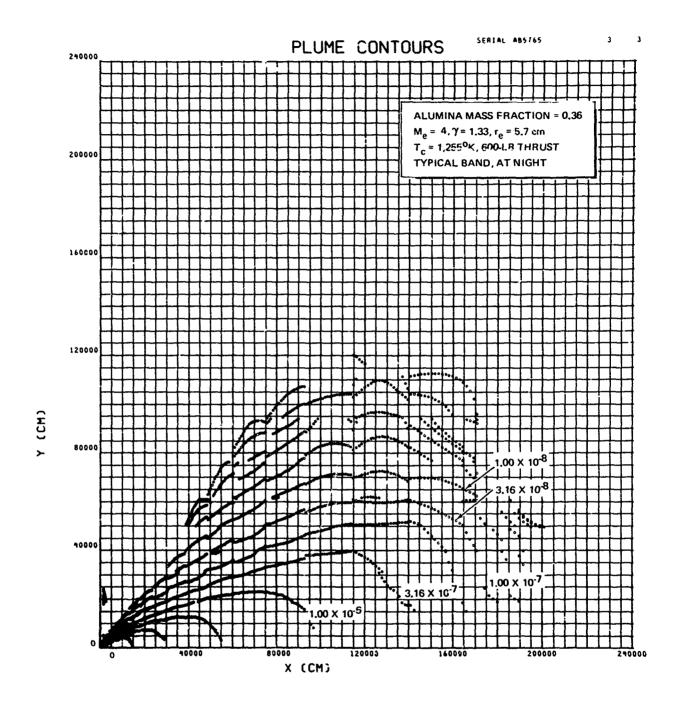


Figure 1-3. Isointensity Lines for Look at 450 from Plume Axis of Solid Propellant Rocket

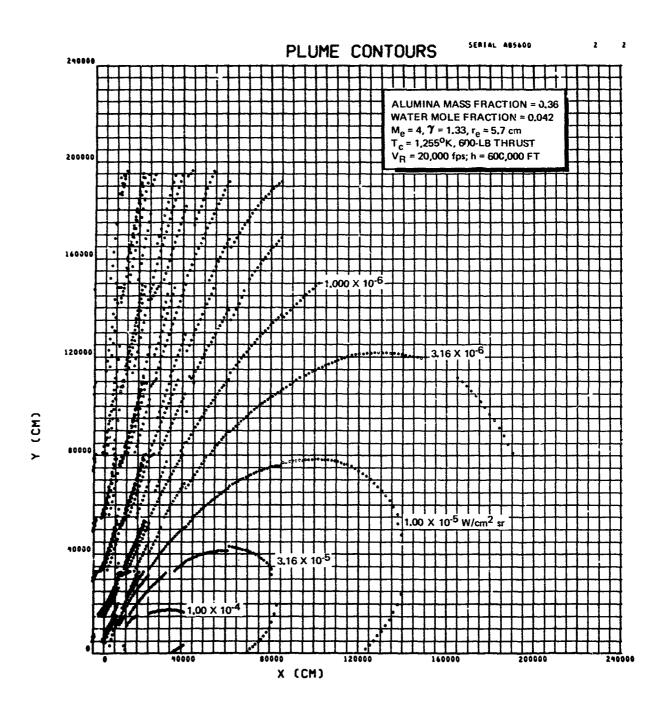


Figure 1-4. Isointensity Lines for Broadside Look at Particle and Excited Water Radiation from Solid Propellant Röcket

Section 2 THEORY

The major elements of the theory used as the basis for the FLAME code is outlined in this section. The approximate model of the far-field multiphase plume is described. For two-phase plumes, closed-form expressions for particle concentration as a function of plume coordinates and particle size are presented. The radiative mechanism for the particle cloud is simply emission, with an effective minimum temperature defined by solar illumination or scattering of Earthshine; it need not be elaborated on here. Gas radiation is considered to be caused by excitation of water rotational spectrum by collision with the atmosphere; this theory is outlined in subsection 2.3. The model can be generalized to include other radiating rotors.

The approaches used were based on existing theoretical work. No attempt has been made to derive new theoretical expressions or include radiative mechanisms not yet understood in the current FLAME. The program is flexible enough, however, to allow the inclusion of other mechanisms at some future time.

2.1 PLUME GAS FLOW

Closed-form approximation to the plume gas is relied upon for flow applicable to the far-field plume. Atmospheric mixing and the resulting plume slow down have not been addressed. Atmospheric excitation of the unaltered plume is considered.

Closed-form solutions have been developed for the density profiles as a function of engine operating conditions. It appears that all properties of the far-field gaseous plume can be thus treated. The computer program will respond to input of engine thrust (or chamber pressure), chamber temperature, expansion ratio (ϵ) , γ (or exit mach number, M_e), and nozzle lip angle by generating a complete history of the rocket plume.

To predict transition to free-molecular-flow, a relationship was derived to express the Knudsen number (K_n) variation in a bipropellant rocket engine plume as a function of the polar coordinates (r, θ) and basic engine parameters F (thrust), C_F (thrust coefficient), T_c (combustion temperature), M_e (exit Mach number) and (ratio of specific heats). The gas viscosity, μ , was assumed to vary as $T^{1/2}$. The engine chamber pressure, P_c , could be introduced through the basic thrust equation:

$$F = P_c A_t C_F$$
where:
$$A_t = Throat Area$$

These derivations used Roberts' approximation (Reference 1) for the plume gas density which indicated that the density was inversely proportional to r^2 and proportional to $\cos\theta$ raised to the power Y(Y - 1) M_e^2 -2.

The derived relationship is shown in Figure 2-1 as a function of M_e , γ , and ϵ (engine expansion area ratio). This information may be employed to determine the shape of the surface (r,θ) required to achieve a given K_n . It is noted that r increases with increasing γ for specified values of T_c , K_n , F, and θ . Indeed, this relationship indicates that $r \to 0$ as $\theta \to 90$ degrees, i.e., that transition occurs at the nozzle lip where, theoretically, there is an infinite pressure gradient as the flow undergoes a Prandtl-Meyer expansion.

Figure 2-2 presents the variation of plume density on a spherical cap (r = constant) as a function of θ and M_e with Y = 1.4. As shown, the density has been normalized by its value on the engine centerline, i.e., θ = 0. When M_e = 2.312, the density distribution is linear with cos θ and becomes increasingly dependent on cos θ as M_e increases, e.g., for M_e = 5.0, ρ ~ (cos θ)¹².

2.2 PARTICLES

A combined nozzle and plume flowfield calculation was performed for an aluminized solid propellant motor to obtain correlations of particle characteristics in the near field. Closed-form solutions, analogous to those obtained for gas-phase-only plumes were not possible due to the nonequilibrium nature of the expansion involving momentum (velocity) and energy (temperature) differences between the phases.



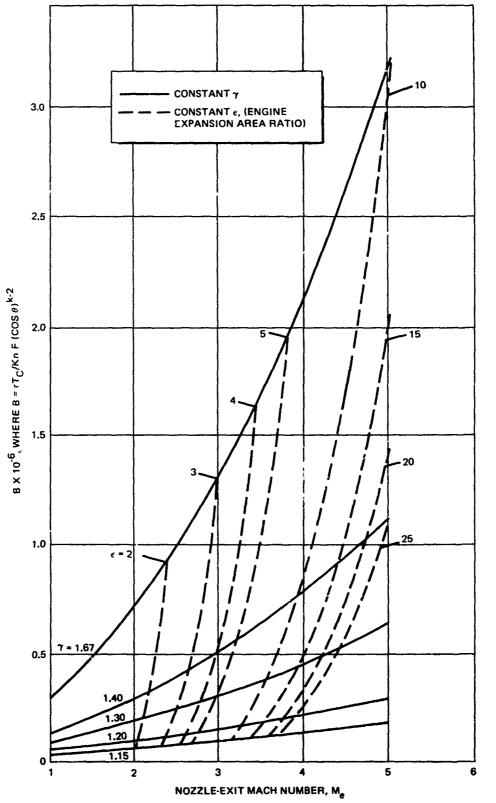


Figure 2-1. Knudsen Number Variation with Engine Conditions

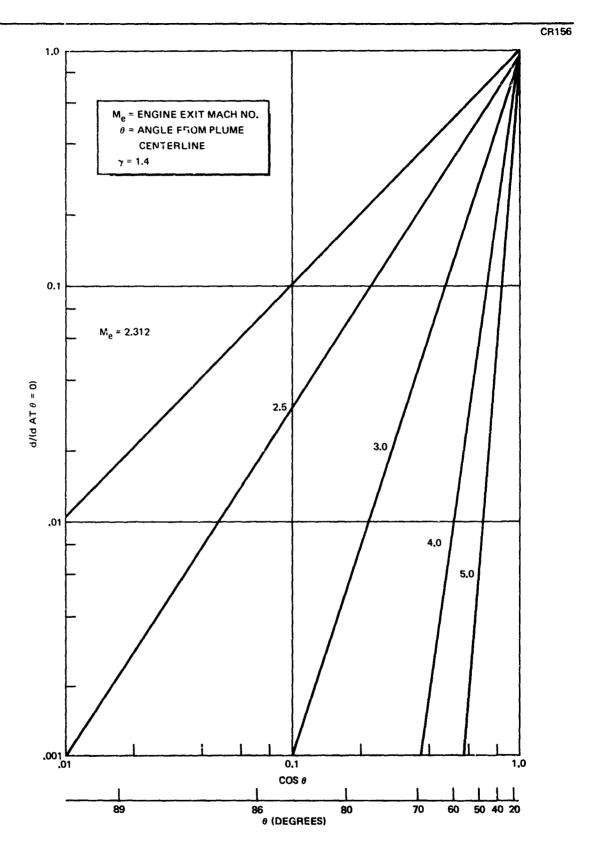


Figure 2-2. Plume Density Distribution on Spherical Cap (r = constant)

Correlations were obtained which related particle velocity, temperature, density, and limiting streamline direction to the seven particle-size groups considered. An expression was derived which related the particle Knudsen number to plume coordinates and motor operating conditions. The primary use of this expression was to determine the extent of the particle continuum region.

Initial estimates of particle emissivity were made, based primarily on existing calculations for alumina. The details of these analyses are presented in later subsections.

2.2.1 Two-Phase Flowfield Particle Correlations

The computer programs described in References 2 and 3 were employed to calculate the nozzle and plume flowfields for an aluminized solid-propellant motor with the following characteristics:

Thrust = 15,000 lb_f

Chamber pressure = 550 psia

Expansion area ratio = 30.8

Aluminum weight content = 16 percent

A total of seven aluminum-oxide particle groups were utilized with particle radii, ρ_i , ranging from 0.6 to 4.7 μm . The corresponding number count versus ρ_i was obtained from Reference 4.

The computed output contained particle velocity, temperature, flow directions, and nondimensional number density for each group at flowfield mesh points within the plume. This information was processed to extract certain particle asymptotic characteristics, namely; velocity, temperature, and flow direction along the particle-limiting streamlines. (The limiting streamline is defined as the boundary outside of which a given particle size will not be present.) It was found that these streamlines became straight lines after the particles had traversed a relatively short distance due to the rapid decay of interphase momentum and energy exchange.

The values of particle temperature, T_i , flow angle, θ_i , and velocity, $V(P_i,\theta)$, on the limiting streamlines are shown in Figure 2-3 as a function of P_i . It is noted that particles larger than 4μ are still undergoing the liquid-solid phase transition at the aluminum-oxide melting temperature of 4,170 R and that these particles are contained within a 23.5-degree halt-angle cone. Only forced convection heat transfer between the particles and the gas is considered in the computer programs and, therefore, a further cooling by radiation must be considered separately.

The plume computer output was employed to obtain the particle velocities for the seven discrete-size groups considered. The correlation was obtained in the form:

$$V(\rho_i, \theta) = [a_i + b_i \exp(-c_i R/r_t)] (\cos \theta)^n i, 0 \le \theta \le \theta_{\rho_i}$$
 (2-1)

where: R, θ = Polar coordinates in plume

r, = Nozzle throat radius

 θ_{ρ_i} = Particle limiting streamline inclination (Reference 1)

Table 4-1 in Section 4 shows the correlation coefficients. The a_i coefficients decrease with increasing particle diameter which indicates that the larger particles achieve a lower limiting velocity as $R/r_t >> 1$. In addition, the n_i coefficients were all negative which indicates that the particle velocities increase off-axis on a given spherical cap. This effect is due to the higher particle concentration on-axis which has caused more interphase momentum transfer from the gas to the particles, i.e., on a per particle basis the gas is less about to "drag" the particle. The lower off-axis particle density allows the gas to "drag" the particles to higher velocities.

The particle density correlation was redone specifically for the off-axis dependency:

$$\frac{\mathrm{d}\rho_{i}}{\mathrm{d}\rho_{o}} = a_{i} \left(\frac{r_{t}}{R}\right)^{n_{i}} \left(\cos\theta\right)^{b_{i}}, \ 0 \le \theta \le \theta_{\rho_{i}}$$
(2-2)

The correlation coefficients are given in subsection 4.5.1.4.

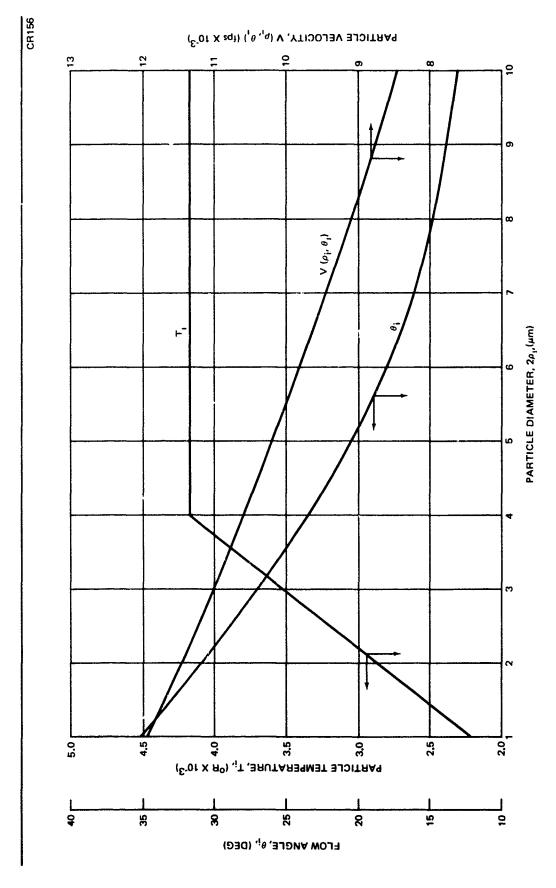


Figure 2-3. Particle Temperature, Flow Angle, and Velocity on Limiting Streamline

Aside for the largest particle size considered, the correlation showed that $b_i > 0$, i.e., the particle density decreased off-axis.

The two-phase computer results were analyzed to obtain a correlation of the gas phase mass flux on the engine centerline with the result being:

$$\begin{pmatrix} V \frac{d}{d_c} \end{pmatrix} = a \left(\frac{R}{r_t}\right)^{-2.7443}$$

$$\theta = 0$$
(2-3)

where d_c = chamber density

The above result indicates that the mass flux is decaying more rapidly than inverse square with distance for the conditions employed. Additional computer runs should be performed for lower value of the solids loading in the exhaust products to determine if the gas phase mass flux will approach a $\left(\frac{R}{r_t}\right)^{-2} \text{ dependency.}$

Following Reference 5, the off-axis gas phase mass flux was assumed to be:

and θ_{α} = Nozzle lip angle at exit

vmax = Maximum Prandtl-Meyer angle

 v_e = Prandtl-Meyer angle at exit Mach number

An expression was derived which relates the particle Knudsen number, Kn_{p_i} to engine operating conditions, location, and particle size for a two-phase plume. This expression was obtained by defining Kn_{p_i} as the ratio of the relative Mach number and Reynolds number, i.e.,:

$$Kn_{p_i} = \frac{Mr_i}{Re_i} = \frac{\left|V - V_{p_i}\right| / \sqrt{g_c \gamma RT}}{d \left|V - V_{p_i}\right| 2\rho_i / \mu}$$
(2-5)

where:

V = Gas velocity

V_{p;} = Velocity of the ith particle

 $g_c = \text{Conversion factor} (= 32.2 [lb_m/lb_f]). (ft/sec^2)$

Y = Gas ratio of specific heats

R = Gas constant

T = Gas static temperature

d = Gas density

 ρ_i = Radius of the ith particle

μ = Gas viscosity

Roberts' solution (Reference 1) was employed to relate d to location in the plume and to certain engine operating parameters; the viscosity was assumed to vary as \sqrt{T} ; and the isentropic relation was used to eliminate T. With these substitutions Equation 1 becomes:

$$K_{p_{i}} = \left[\frac{12x46.6x10^{-10} \sqrt{R_{o}/g_{c}} T_{c}}{P_{c}^{\gamma 3/2} (\dot{\gamma}-1) M_{e}^{2} \left(1+\frac{\gamma-1}{2} M_{e}^{2}\right)} - \frac{1}{\gamma-1} \right] \frac{(r/r_{e})^{2}}{\rho_{i} \cos^{-k-2}\theta}$$

where:

r = Nozzle exit radius

R_o = Universal gas constant (=1545 ft. lb_f/mole ^oR)

T_c = Combustion temperature

P_c = Chamber pressure

M_e = Nozzle exit mach number

 r, θ = Polar coordinates of point in plume

k = Hypersonic similarity parameter (= $Y(Y-1)m^2$)

The above equation indicates that, for given engine operating conditions and plume coordinates, the smaller particles will achieve the higher Knudsen number, i.e., they will "transition" sooner. Numerical calculations for a l μ radius particle and typical engine operating conditions indicates that Kn reaches unity very close to the nozzle exit plane. By comparison a gasphase plume achieves a unity Knudsen number only at large distances from the exit plane.

2.2.2 First Estimate of Emissivity of Alumina Particles

The Mie theory for scattering of electromagnetic waves by dielectric spheres forms the basis for theoretical determination of particle emissivity. The theory actually computes the scattering and absorption efficiency factors, $Q_s(\alpha)$ and $Q_a(\alpha)$, defined by

$$Q = \sigma/\sigma_g$$

where σ_g is the geometrical cross section and σ the actual cross section. α is the size parameter $2\pi a/\lambda$, where a is the particle radius. The values obtained for Q are determined by the values of the real and imaginary parts of the index of refraction n, related by

$$n = n_1 - in_2$$

Since $n = n(\lambda, T)$, a more accurate functional dependence of Q would be $Q=Q(\lambda, a, T)$.

Plass (References 6 and 7) has computed efficiency factors for alumina spheres of 0.1 to 10μ for a range of wavelengths 0.5 to 10μ . Values of n_1 and n_2 used in this wavelength range were based upon measurements made with large sapphire crystals. They indicates $n_2/n_1 \ll 1$.

For determination of the emissivity, ϵ , Plass used an equation originally derived for the emissivity of a semi-infinite homogenous slab, with

$$\epsilon = 2.3 \sqrt{\sigma_a/\sigma_s}$$

Thus the following values of ϵ are obtained from the following table:

λ	r	€
2μ	1.1μ	0.005
	5.1	0.006
	9.9	0.018
5μ	1.1	0.054
	5.1	0.089
	9.9	0.115

(Figures for $\lambda = 10\mu$ are not immediately available).

It is to be noted that in the range $10 < \lambda < 20\mu$, values of n_2 increase by more than an order of magnitude (Brannon and Goldstein, Reference 8) so ϵ would be expected to be larger there, aside from size effects.

The results of Plass have been questioned by Morizumi and Carpenter (Reference 9) on the grounds that measurements of plume emissivities of rocket exhaust indicate larger values for particle emissivities. This conclusion is rationalized on the grounds that absorbing properties of rocket alumina particles are higher than those of sapphire, due to the polycrystal-line properties of the former. They conclude that emissivity of alumina particles in rocket exhausts lies between 0.1 and 0.3.

It is concluded that conflicting and insufficient data available at this time precludes a definitive answer for the emissivity in the infrared. For provisional values, the following are suggested, using linear interpolation and extrapolation to obtain values at other sizes or wavelengths.

<u>λ</u>	<u>r</u>	€	_λ	<u>r</u>	_ €	_λ	<u>r</u>	_€
5μ	lμ	0.05	10μ	1μ	0.1	20μ	$l\mu$	0.2
	5	0.09		5μ	0.15		5	0.25
	10	0.12		10μ	0,20		10	0.30

An approximate analytic correlation of the above data for Al₂0₃ emissivity was found. The correlation is shown in Figure 2-4, and is adequate, considering the coarseness of the available data. The emissivity correlation is of the form

$$\epsilon_{\lambda}(\mathbf{r}) = (\lambda + 2.5)(.00689 + 11.32 \,\rho - 2220 \,\rho^2)$$
 (2-6)

for λ in microns and ρ (particle radius) in cm.

2.3 ATMOSPHERIC EXCITATION

Atmospheric excitation of rocket-plume species by collisional high relative velocities has been identified as a strong source of plume radiation. A detailed analysis of the problem was presented in Reference 10 and summarized in Reference 11. The unique feature of this analysis is the inclusion of the time-dependent radiative decay in the treatment of the excited species. Thus, there is a finite time required before the plume can generate radiation by this mechanism.

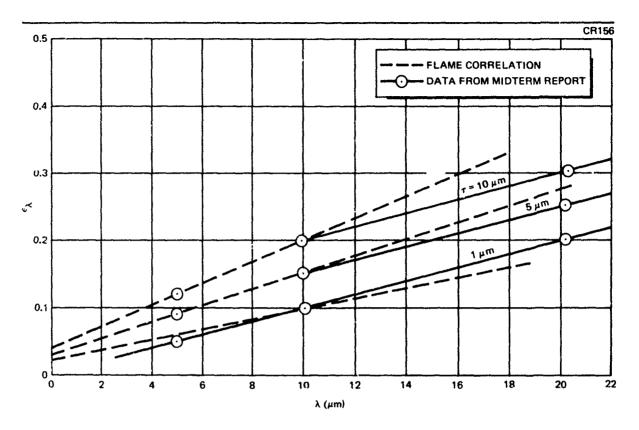


Figure 2-4. Alumina Particle Emissivity

Detailed calculations of trajectories of the colliding species, averaged over initial orientations, were done using Program P2160. Figure 2-5 shows the resulting rotational temperatures induced as a function of impact parameter and V_{∞} . Also shown on Figure 2-5 is the linear variation found for $T_{\rm m}/V_{\infty}$ as a function of V_{∞} ($T_{\rm m}$ is the maximum rotational temperature for any impact parameter).

The effective collision cross section (S) to use with $T_{\bf m}$ is found by requiring $T_{\bf m}S=2\,\pi\, {\int_0^\infty}\, T(b)\ bdb$

Calculated values of S are shown ir Figure 2-6. For the current version of FLAME, S will be assumed to be constant (=0.7x10⁻⁵ cm²).

The atmospheric interaction model used in FLAME follows that of Reference 10 namely:

- A. For the wavelength band of interest, the total energy (E_T) and radiance (j) per molecule are calculated as a function of T.
- B. Radiative decay times are calculated as a function of T_m.



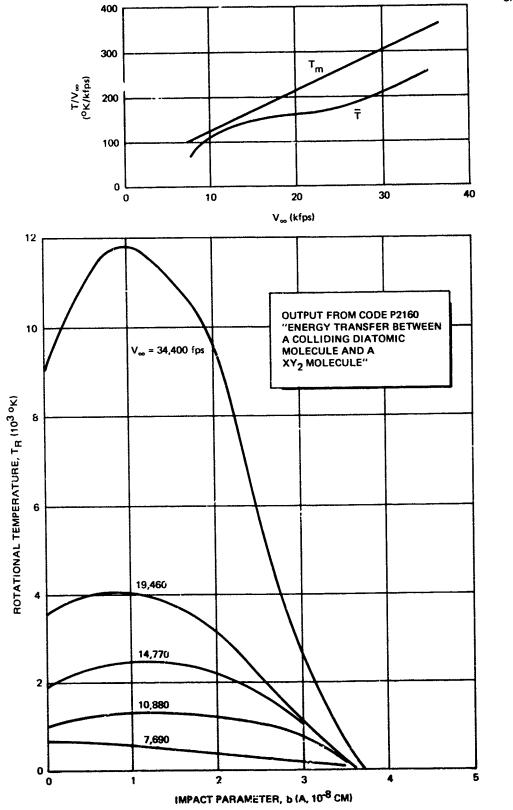


Figure 2-5. Water Rotational Energy from Single Collision

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Figure 2-6. Rotational Excitation Parameters for ${\rm H_2O}$

- C. The time of flight to a position in the plume is compared to the decay time so that an average radiance per excited molecule can be found.
- D. The fraction of molecules that are excited are calculated assuming initially cold plume species.
- E. No mechanism currently is programmed to account for plume/ atmosphere mixing, but plans are to add this later. Currently if a molecule undergoes multiple collisions in the atmosphere, the resultant radiance is found by simple addition.
- F. The radiance decay of an excited species is assumed to be linear in time.
- G. Data currently in program is applicable to a single species, H₂0.

Further details of the analysis are available in Reference 10.

The analysis of Reference 10 was changed in one way. Since the radiative model used assumes single collisions, and the plume may be viewed from long distance, an attenuation factor (A_t) must be introduced to account for depletion of the supply of uncollided molecules as the gas expands far from the nozzle. Assuming each collision removes that molecule from further consideration, the reduction of the number of available molecules from N_0 (initial value) to N (at point r in plume) is given by

$$\frac{N}{N_0} = \exp(-A_t) \tag{2-7}$$

where

$$A_{t} = S \int_{0}^{r} n_{\infty} ds,$$

where S is the collision cross section and n_{∞} is the atmospheric density.

A correlation was obtained for the spectral linear absorption coefficient of the $H_2\,0$ rotational spectrum. The results are that

$$k_{\lambda} = \log^{-1}(y) \quad \text{in cm}^{-1}$$
 (2-8)

where

$$y = A \frac{(\omega - \omega_0)^2}{\omega} + y_0$$

$$\omega_0 = 10.6 \sqrt{T}$$
, T in $^{\circ}$ K

$$y_0 = \frac{279.96}{\sqrt{T} + 50.031} - 3.208$$

$$A = \begin{bmatrix} 1.030 - \frac{105.582}{\sqrt{T} - 2.896} \end{bmatrix} \times 10^{-4}$$

$$\omega = 10,000/\lambda, \lambda \text{ in } \mu\text{m}$$

The data correlated are shown in Figure 2-7, where the results of the correlation are shown for temperatures of 600°K and 1,800°K.

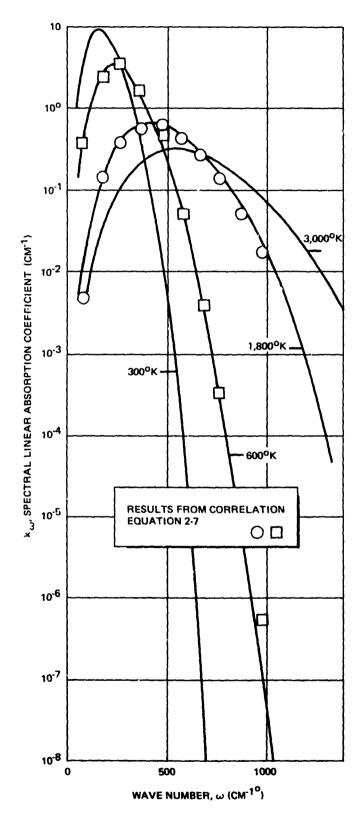


Figure 2-7. Absorption Coefficient from Water Rotation at One Atmosphere Pressure

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Section 3 FLAME CODE

The following subsections deal with the logic underlying the FLAME code as it is now configured. Subsection 3.1 describes the logic that went into the choice of coordinate systems, and displays the geometric model used. Subsection 3.2 presents the flow charts used as the basis for FLAME.

3.1 GEOMETRY CONSIDERATIONS

The rocket plume is an axisymmetric body about the z-axis, as illustrated in Figure 3-1. Integration through the plume is carried out through the conical volume in the first and second quadrants defined by the polar coordinates θ_m and r_m . The viewing angle is specified completely by the angle α with respect to the x-axis in the xz plane.

We want to integrate emittance along families of lines of sight, such as s,

$$\int E(\mathbf{r}, \boldsymbol{\theta}) ds$$

We can express ds in terms of either cartesian or spherical coordinates,

$$ds = dx \sqrt{1 + \left(\frac{dy}{dx}\right)^2 + \left(\frac{dz}{dx}\right)^2}$$
 (3-1)

or

$$ds = dr \sqrt{1 + r^2 \left(\frac{d \phi}{dr}\right)^2 + r^2 \sin^2 \phi \left(\frac{d\theta}{ar}\right)^2}$$
 (3-2)

Since emittance is a function of r, θ along, equation (3-2) is more inviting to use, if it is practical. The following analysis shows, however, that cartesian coordinates will probably result in less computation time, and should be used.

The conical segment in Figure 3-1 has equation

$$x^2 + y^2 = z^2 tan^2 \theta_1$$
 (3-3)

The family of lines along direction of observation, α have equations

$$y = y_1 \tag{3-4}$$

$$x = x_1 + s \cos \alpha \tag{3-5}$$

$$z = z_1 + s \sin \alpha \tag{3-6}$$

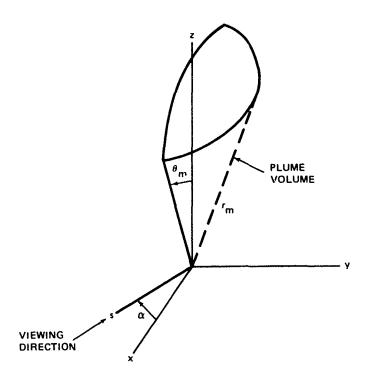


Figure 3-1. Plume Orientation and Viewing Direction

These equations can be switched to spherical coordinates using the standard relations

$$x = r \sin\theta \cos\phi \tag{3-7}$$

$$y = r \sin \theta \sin \phi$$
 (3-8)

$$z = r\cos\theta \tag{3-9}$$

The following discussion derives the spherical and cartesian forms for ds along straight lines of sight:

3.1.1 Spherical Coordinates

From Equation 3-8
$$\frac{y_1^2}{(r\sin\theta)^2} = \sin^2\phi;$$
 (3-10)

From Equations 3-5, 3-6, 3-7, and 3-9

$$\frac{\left[\cot \alpha \left(r\cos\theta-z_1\right)+x_1\right]^2}{\left(r\sin\theta\right)^2}=\cos^2\emptyset$$
(3-11)

Adding Equations 3-10 and 3-11

$$y_1^2 + \left[\cot\alpha \left(r\cos\theta - z_1\right) + x_1\right]^2 = \left(r\sin\theta\right)^2 = r^2 - \left(r\cos\theta\right)^2$$
 (3-12)

let

$$\cot \alpha \quad r \cos \theta = F$$

$$-\cot \alpha \quad z_1 + x_1 = A$$

$$-y_1^2 + r^2 = B$$

Equation 3-12 becomes

$$F^2 + 2AF + A^2 - B + \frac{F^2}{\cot \alpha} = 0$$

or

$$F^2(1+\tan^2\alpha) + 2AF + A^2 - B = 0$$
; $F^2/\cos^2\alpha + 2AF + A^2 - B = 0$

Therefore

$$F = \frac{-2A \pm \sqrt{4A^2 - 4A^2 \left(\frac{1}{\cos^2 \alpha}\right) + \frac{4B}{\cos^2 \alpha}}}{1/\cos^2} =$$

$$-2A \cos^2 \alpha \pm 2 \cos \alpha \sqrt{B - A^2 \sin^2 \alpha}$$

Therefore

$$\sin \alpha r \cos \theta = -2A \cos \alpha \pm 2 \sqrt{B-A^2 \sin^2 \alpha}$$

or

$$\cos \theta = -\frac{2 \operatorname{Acot} \alpha}{r} \pm \frac{2}{r} \quad \sqrt{\frac{B}{\sin^2 \alpha} - A^2} \approx u(r)$$

$$\theta = \cos^{-1}\left(\frac{2A\cot\alpha}{r} \pm \frac{2}{r} \sqrt{\frac{B}{\sin^2\alpha}A^2}\right)$$

From consideration of the case $\alpha = 90$ degrees and noting that $\cos \theta > 0$, only + root is valid.

$$\frac{d\theta}{dr} = \frac{-1}{\sqrt{1-u^2}} \frac{du}{dr}$$

or

$$\frac{d\theta}{dr} = \frac{-1}{\sqrt{1-u^2}} \left[-\frac{u}{r} + \frac{1}{r} \left(\frac{B}{\sin^2 \alpha} - A^2 \right)^{-1/2} \frac{2r}{\sin^2 \alpha} \right]$$
 (3-13)

also

$$\sin \phi = \frac{y_1}{r \sin \theta} = \frac{y_1}{r \sqrt{(1-u^2)}} \equiv v(r)$$

$$\frac{d\phi}{dr} = \frac{1}{\sqrt{1-v^2}} \frac{dv}{dr}$$

where

$$\frac{dv}{dr} = \frac{-y_1}{r^2 \sqrt{1-u^2}} + \frac{y_1 u}{r(1-u^2)^{3/2}} \frac{du}{dr}$$

From Equation 3-13

$$\frac{d\theta}{dr} = \frac{uw}{r(1-u^2)} \frac{1}{2}$$

where

$$w(r) = \left[1 - \frac{2r}{u\sin^2\alpha - \sqrt{B/\sin^2\alpha - A^2}}\right]$$

then

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$$\frac{dv}{dr} = \frac{-y_1}{r^2 \sqrt{1-u^2}} - \frac{y_1 u^2 w}{r^2 (1-u^2)^{3/2}}$$

$$\frac{du}{dr} = -\frac{uw}{r}$$

or finally

$$ds = dr \left[1 + r^2 \frac{1}{1 - v^2} \frac{y_1^2}{r^4 (1 - u^2)} \left(1 + \frac{u^2 w}{1 - u^2} \right)^2 + r^2 \frac{y_1^2}{r^2 (1 - u^2)} \frac{u^2}{r^2} w^2 \right]^{1/2}$$
(3-14)

From Equations 3-3, 3-4, 3-5, and 3-6 the limits of integration, r, and r_2 can be found.

We have

$$x^{2} + y^{2} + z^{2} = r^{2}$$
 (3-15)
 $x^{2} + y^{2} = z^{2} \tan^{2} \theta_{m}, = r^{2} - z^{2}$
 $r^{2} = z^{2} (1 + \tan^{2} \theta_{m})$

Therefore

$$z = r \cos \theta_{m}$$

Also

$$y = y_1$$

and

$$\frac{x-x_1}{\cos\alpha} = \frac{z-z_1}{\sin\alpha}$$

or

$$x = \cot \alpha (z - z_1) + x_1 \tag{3-16}$$

Substituting x, y, and z into equation of the cone, Equation 3-3 we have a quadratic equation for r, namely

$$(\cot^{2}\alpha\cos^{2}\theta_{1} + \cos^{2}\theta_{1} - 1)r^{2} + (2x_{1}\cot\alpha\cos\theta_{1} - 2\cot^{2}\alpha z_{1}\cos\theta_{1})r$$

$$+ (\cot^{2}\alpha z_{1}^{2} - 2x_{1}\cot\alpha z_{1} + x_{1}^{2} + y_{1}^{2}) = 0$$
(3-17)

The solution gives r_1 and r_2 , as function of z_1 , x_1 , and y_1

3.1.2 Cartesian Coordinates

From Equation 3-16

$$z = \tan \alpha (x-x_1) + z_1 \tag{3-18}$$

or

$$\frac{\mathrm{dz}}{\mathrm{dx}} = \tan \alpha$$

Therefore

$$ds = dx \sqrt{1 + \tan^2 \alpha} = \frac{dx}{\cos \alpha}$$
 (3-19)

or

$$ds = \frac{dz}{\sin \alpha} \tag{3-20}$$

We obtain r and θ from Equation 3-18

$$r^2 = z^2 + y_1^2 + \left[\cot \alpha (z - z_1) + x_1\right]^2$$
 (3-21)

$$\cos \theta = z/r \tag{3-22}$$

It is apparent that it is easier to use cartesian coordinates, and Equations 3-19 or 3-20 together with Equations 3-18, 3-21, and 3-22 to integrate through the plume, than to use the single, complex Equation 3-14.

To avoid singularities at $\alpha = 0$ degree or 90 degrees, x will be used as the independent variable for $0 \le \alpha \le 45$ degrees and z for the independent variable for 45 degrees < $\alpha \le 90$ degrees. Since the plume is transparent, negative values of α are not needed, the integrals being equal to values for 90 degrees + α .

For cartesian coordinates, the intercept equations are more complex than for spherical coordinates; substituting

$$y = y_1 \tag{3-23}$$

$$x = \cot \alpha (z-z_1) + x_1$$
 (3-24)

into equation of spherical cap

$$x^2 + y^2 + z^2 = r_m^2 ag{3-25}$$

and conical sides

$$x^2 + y^2 = z^2 tan^2 \theta_m$$
 (3-26)

yields the following equations for the resulting four intercept points of a straight line of sight:

For $0 < \alpha < 45$ degrees (choose $x_1 = 0$)

$$(1 + \tan^2 \alpha) x^2 + (2z_1 \tan \alpha) x + y_1^2 + z_1^2 - r_m^2 = 0$$
 (3-27)

$$(\cot^2 \theta_m - \tan^2 \alpha) x^2 - (2z_1 \tan \alpha) x + y_1^2 \cot^2 \theta_m - z_1^2 = 0$$
 (3-28)

and for $45 \text{ degrees} < \alpha \le 90 \text{ degrees}$ (choose $z_1 = 0$)

$$(1-\cot^2\alpha)z^2 + (2x_1\cot\alpha)z + y_1^2 + x_1^2 - r_m^2 = 0$$
 (3-29)

$$(\tan^2 \theta_{\rm m} - \cot^2 \alpha) z^2 - (2x_1 \cot \alpha) z - (y_1^2 + x_1^2) = 0$$
 (3-30)

For each range of α , solution of the two quadratic equations will yield four values for x or z. The correct limits of integration are found by selecting those pairs of x or z for which the corresponding values of r and θ (by Equations 3-18, 3-21, and 3-22) are on the real plume boundary, namely

$$r_{m}$$
 and $\theta \leq \theta_{m}$

or

$$\theta_m$$
 and $r \le r_m$

3.1.3 Integration Schemes and Step Sizes

Most of the integrals of the form

$$I = \int_{\mathbf{x}_1}^{\mathbf{x}_2} f(\mathbf{x}) d\mathbf{x}$$
 (3-31)

are evaluated using Gaussian eight-point quadrature. Defining

$$y = (2x - x_1 - x_2)/(x_2 - x_1)$$
 (3-32)

we have

$$I = \frac{x_2 - x_1}{2} \int_{-1}^{1} f(x) dy$$
 (3-33)

Then

$$I \approx \frac{x_2 - x_1}{2} \sum_{i=1}^{8} w_i f(x_i)$$
 (3-34)

where

$$y_1 = y_8 = 0.18343 46424 95650$$

$$y_2 = y_7 = 0.52553 24099 16329$$

$$y_3 = y_6 = 0.79666 64774 13627$$

$$y_4 = y_5 = 0.960289856497536$$

and

$$w_1 = w_8 = 0.36268 37833 78362$$

 $w_2 = w_7 = 0.31370 66458 77887$
 $w_3 = w_6 = 0.22238 10344 53374$
 $w_4 = w_5 = 0.10122 85362 90376$

The values of x_i are related to y_i through Equation 3-32.

For integrating through the plume, Simpson's rule is used

$$\int_{\mathbf{x}}^{\mathbf{x}+2\Delta\mathbf{x}} f(\mathbf{y}) d\mathbf{y} = (f(\mathbf{x}) + 4f(\mathbf{x}+\Delta\mathbf{x}) + f(\mathbf{x}+2\Delta\mathbf{x})) \frac{2\Delta\mathbf{x}}{3}$$
 (3-35)

The step size, Δx , is varied after each successive integration from x to $x + 2\Delta x$. A geometric variation of Δx assures that the smallest step size occurs as the integration along the line of sight passes through the plume center (x = 0 or z = 0 planes).

Let the total maximum distance along the line of sight be \mathbf{x}_T . Starting at the middle of the plume and integrating out, a distance $\mathbf{x}_T/2$ is covered in about n/2 steps (n is a program input). If each $\Delta \mathbf{x}$ step is r times bigger than the previous, the rule for geometric progression gives

$$\frac{x_{T}}{2} = \Delta x_{\min} \frac{r^{n/2} - 1}{r - 1}$$
 (3-36)

where Δx_{min} is the initial, smallest step size. Choosing a 10-percent increase in Δx each step, r = 1.1

or

$$\Delta x_{\min} = \frac{x_{\text{T}}}{20(1.1^{n/2}-1)} \tag{3-37}$$

Starting at position $x = -|x_0|$, a large step size is used which decreases until x = 0 and then increases again. The initial value of Δx is found from the formula for the sum in terms of the last term (Δx_{in})

$$x_{o} = \frac{\Delta x_{in} r - \Delta x_{min}}{r - 1}$$
 (3-38)

or

$$\Delta x_{in} = (0.1 x_{c} + \Delta x_{piin})/1.1$$
 (3-39)

A similar stepping scheme is used to obtain a higher density of lines of sight near the rocket exit plane or plume centerline than used in the rarified plume.

3.1.4 Relative Velocity

The rocket-body velocity must be specified in the coordinate system of Figure 3-1 so that relative velocity of plume gases and atmosphere can be calculated. Figure 3-2 shows the position of the vehicle velocity vector V_R , and the definition of straight up, \mathring{u} .

Integration is performed only over a plume where y > 0. To obtain other half of plume, Ψ_R and Ψ_u should be increased 180°.

Relative velocity is just the vector sum of $V_p \hat{r}$ and V_R where \hat{r} is the unit vector along the polar coordinate r to a point in the plume being considered. Altitude at this same point is missile altitude h_0 plus projection of \vec{r} onto \hat{u} .

$$h = h_0 + \vec{r} \cdot \hat{u}$$
 (3-40)

3.2 FLOW CHARTS

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This section presents the logic flow of the FLAME code. Input data are listed in Tables 3-1 through 3-4. The Figure 3-3 is a simplified diagram of the overall code. The charts (Figures 3-4 through 3-9) expand upon certain regions of the computer.

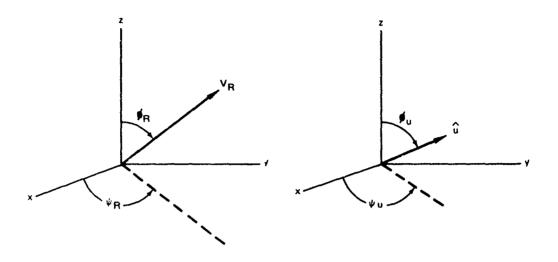


Figure 3-2. Location of Vehicle Velocity Vector and Upward Direction in Coordinate System of Figure 3-1.

Table 3-1

LISTING OF INPUT DATA

	DIDITIO OF HIT OF BILLIE		Real or
Read in		FORTRAN	
t,	time after ignition (sec)	TIG	R
r _e ,	nozzle exit radius (cm)	RE	R
γ,	ratio of specific heats (>1)	GAMMA	R
М _е ,	exit mach number	ME	R
θ _e ,	nozzle lip angle (deg)	LIP	R
d _e (t),	exit plane density as a table versus time after ignition (g/cm ³ vs sec)	PDE TID	R R
I _{ND} ,	night/day flag (0 or 1)	IND	I
I _{SL} ,	solid/liquid flag 0 or -2 (gas); 1 or -1 (solid)	ISL	I
$\lambda_{\min,j}$		WLMIN	R
λ_{\max}	band limits (μm)	WLMAX	R
n,	approximate integration steps through layer	INT	Ι
m,	approximate number of lines of sight used in one direction	LINES	I
T_c	chamber temperature (OK)	TC	R
×m,	radiating gaseous species mole fraction*	XM	R
$^{\mathrm{T}}\mathrm{E}^{'}$	Earth effective temperature (OK)	TE	R
α,	look angle ($0 \le \alpha \le 90$ degrees)	AL	R
h ₀ ,	altitude of engine (ft) (from 0.3 Mft to 3 Mft)	НО	R
v_R ,	rocket velocity (fps)	VR	R
$\phi_{\mathbf{u}}$	polar angle of vertical direction (deg)	PHIU	R
$^{\psi}\mathrm{u}$	azimuthal angle of vertical direction (deg)	XIU	R
$\phi_{\mathbf{R}}$,	polar angle of rocket-body velocity (deg)	PHIR	R
$\Psi_{\mathbf{R}}$,	azimuthal angle of rocket-body velocity (deg)	XIR	R
END,	flag signifying end of run (0 or 1)	END	I

^{*}FLAME assumes an average molecular weight (WMOL) of 17 for the plume species (approximately correct for pure gas plumes, or the gaseous constituents of solid-propellant plumes). When gaseous radiation from a solid propellant plume is calculated ($|I_{SL}| = 1$ and $x_m \neq 0$), the molecular weight should be corrected to include solids. This is done by modifying x_m , since only the ratio XM/WMOL is used by FLAME. For a typical solid-propellant plume, WMOL = 26, so XM should be reduced by the ratio (17/26).

Table 3-2

PROGRAM REGION 1. INITIAL CALCULATIONS-PARTICLES

The following data will be calculated for each particle radius, p_i, corresponding to the ith particle size:

EPT,	table of total emissivity versus temperature
TCOOL,	table of particle temperature versus time after solidification
BPASS,	black body intensity in band of interest
ETL	table of average emissivity in region from $^{\lambda}\min_{\text{to }^{\lambda}\max}$ versus temperature
TMIN,	equilibrium temperature of particle
DELT,	time required to solidify droplet

Table 3-3

PROGRAM REGION 2. INITIAL CALCULATIONS-GAS

Calculate limiting plume-gas velocity, VP. For the band width specified, calculate

	Tot the sense water epochton, caronicol
ET,	table of total available rotational energy as function of temperature
JM,	table of radiative rate (radiance per molecule) as function of temperature
JT,	table of radiative rate as a function of time

Table 3-4

PROGRAM REGION 3. SELECT x_1 OR z_1 for $a \ge 90 - \theta_m$

Test IA to determine magnitude of a. For IA = 0, define z_1 . For IA = 1, define x_1 . Both x_1 and z_1 are labeled X1 in the program.

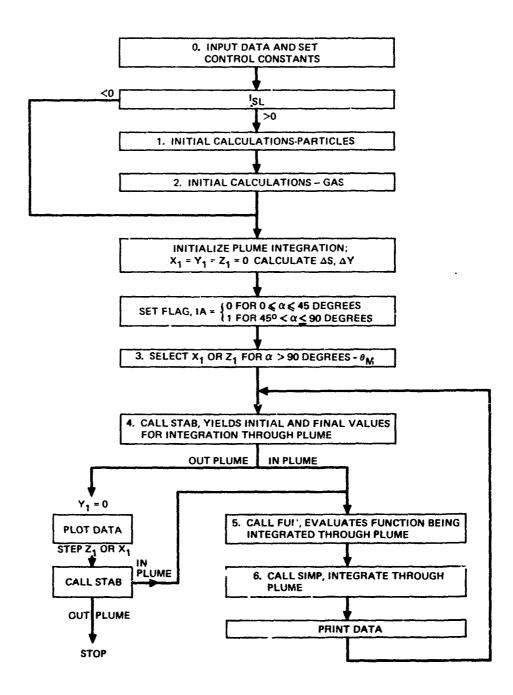


Figure 3-3. Master Flow for Program P2170, FLAME

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CALL STAB AND CHECK (YIELDS INITIAL AND FINAL VALUES FOR INTEGRATION THROUGH PLUME)

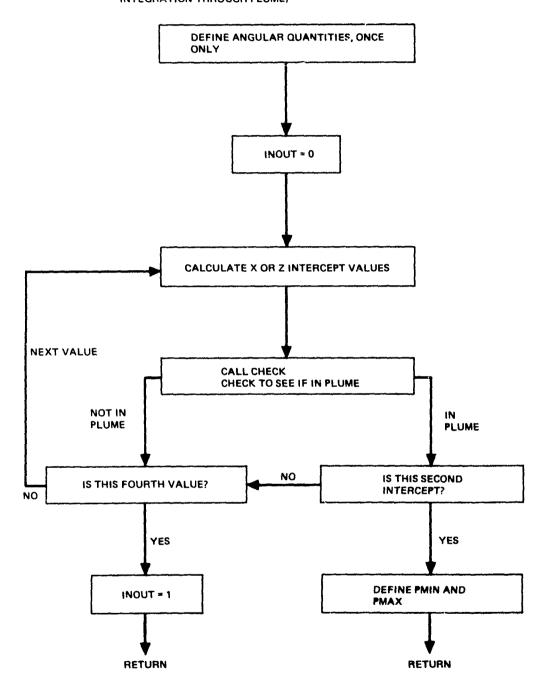


Figure 3-4. Program Region 4 Flow

CALL FUN (EVALUATES FUNCTION BEING INTEGRATED THROUGH PLUME)

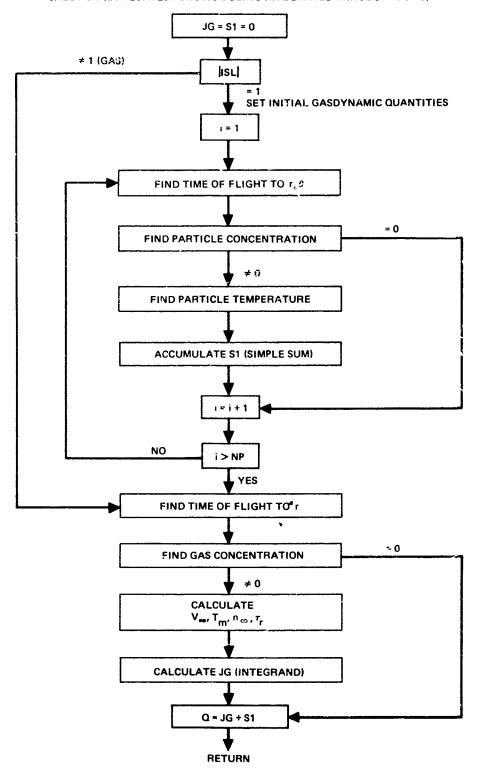


Figure 3-5. Program Region 5 Flow

CAL SIMP (INTEGRATE THROUGH PLUME)

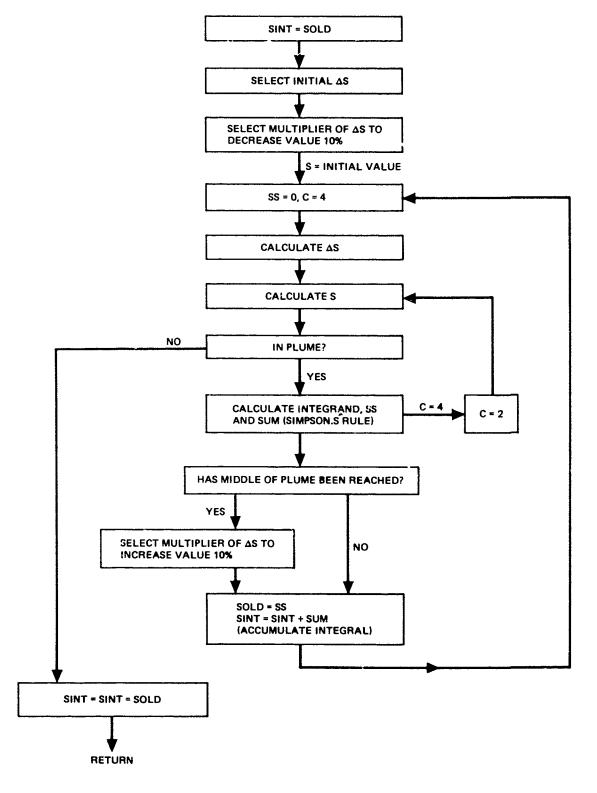


Figure 3-6. Program Region 6 Flow

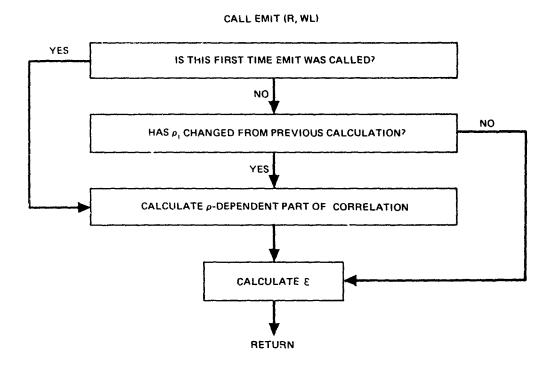


Figure 3-7. Particle Emissivity Routine Flow Chart

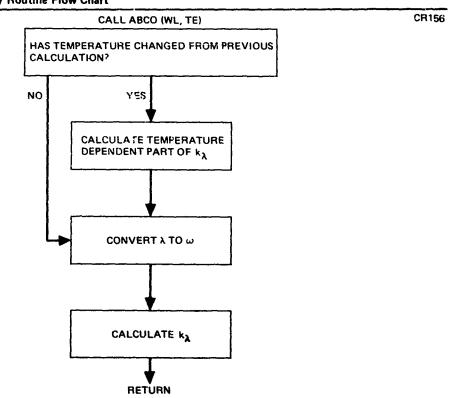


Figure 3-8. Absorption Coefficient Routine Flow Chart

CALL CONTOUR (KP)

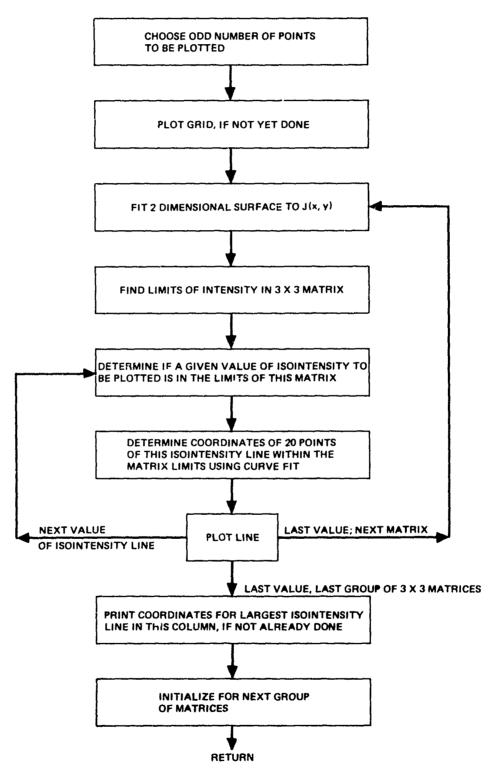


Figure 3-9. Contour Plotter Flow Chart

Section 4 EQUATIONS TO BE SOLVED

The following subsections relate to the program regions shown on the master flow chart, Figure 3-3.

4.1 PROGRAM SECTION 1. INITIAL CALCULATIONS-PARTICLES

4.1.1 EPS, Particle Emissivity Versus Temperature

$$\epsilon_{i}$$
 (T) = $\int_{0}^{\infty} \epsilon_{i\lambda} R_{\lambda}^{0}$ (T) $d\lambda/\sigma T^{4}$ (4-1)

where

$$R_{\lambda}^{O}(T) = \frac{2\pi c^{2}h}{\lambda^{5}} \left[e^{\frac{hc}{\lambda kt}} - 1\right]^{-1} = \frac{c_{1}}{\lambda^{5}} \left[e^{\frac{c_{2}}{\lambda T}} - 1\right]^{-1}$$

$$(4-2)$$

c = velocity of light

σ = Stefan-Boltzmann constant

= $5.672 \times 10^{-12} \text{ j/cm}^2 \text{ oK}^4 \text{ sec}$

 $\epsilon_{i\lambda}$ = correlation eq (2-6)

 $c_1 = 3.742 \times 10^4 \text{ w } (\mu\text{m})^4/\text{cm}^2$

 $c_2 = 14390 \, \mu m^{0} K$

See Subsection 3.1.3 for description of Gaussian Integration.

4.1.2 TCOOL, Particle Temperature Versus Time After Solidification

$$\Delta t_{i} (T) = \frac{c_{p} \rho_{i}}{3\sigma} \int_{T_{f}}^{T} \frac{dT}{\epsilon_{i} T^{4}}$$
(4-3)

where

cp = particle specific heat

Tf = fusion temperature

4.1.3 BPASS Black Body Intensity in Band of Interest

$$B_{\Delta\lambda}^{O}(T) = \frac{1}{\pi} \int_{\lambda_{\min}}^{\lambda_{\max}} R_{\lambda}^{O}(T) d\lambda$$
 (4-4)

4.1.4 EPSI, Average Particle Emissivity

$$\epsilon_{i,\Delta\lambda} (T) = \frac{1}{\pi} \int_{\lambda_{\min}}^{\lambda_{\max}} \epsilon_{i\lambda} R_{\lambda}^{o} (T) d\lambda / B_{\Delta\lambda}^{o} (T)$$
 (4-5)

4.1.5 TMIN, Particle Equilibrium Temperature at day,

$$\epsilon_i T_{\min,i}^4 = \frac{\alpha_i \Gamma}{2\sigma} + \frac{\alpha_i!}{2} T_E^4 \equiv K$$
 (4-6)

where

 $a_i = \text{solar absorptivity} [\sim \epsilon_i (6000^\circ)]$

 a_i^i absorptivity to Earthshine $[\sim \epsilon_i^i (T_E^i)]$

 Γ = solar constant = 0.1353 w/cm²

T_E = effective Earth temperature

at night, Equation 4-6 is used with

 $\Gamma = 0$

Newton-Rapheson solution is of form

$$T_{j} = \frac{1}{3} \left[2T_{j-1} + \frac{K}{T_{j-1}^{3} \epsilon_{i} (T_{j-1})} \right]$$
 (4-7)

4.1.6 DELT, Time to Solidify

$$\Delta t_{i,s} = \frac{\rho_i m \lambda}{3 \epsilon_i \sigma T_f}$$
 (4-8)

where

m = particle density

 λ = particle heat of fusion

4.2 PROGRAM SECTION 2. INITIAL CALCULATIONS -- GAS

4.2.1 Plume Gas Limiting Velocity

$$V_{p} = a_{o} \left(\frac{2}{\gamma - 1}\right)^{1/2} \tag{4-9}$$

where

$$a_{O} = \sqrt{\gamma R T_{C}}$$
 (4-10)

4.2.2 ET, Rotational Energy Versus Temperature

$$E_{T} = \frac{hcB}{2\sigma^{3/2}} \left[x_{1}^{1/2} e^{-x_{1}} - x_{2}^{1/2} e^{-x_{2}} + \frac{\sqrt{\pi}}{2} (erf \sqrt{x_{2}} - erf \sqrt{x_{1}}) \right]$$
(4-11)

where

B = molecular rotational constant

$$\sigma = hcB/kT \tag{4-12}$$

$$x_{1,2} = J_{1,2}^{2} \sigma$$
 (4-13)

$$J_{1,2} = \frac{1}{2B\lambda_{\min, \max}}$$
 (4-14)

$$\frac{hc}{2} = 0.9928 \text{ w sec cm}$$

$$\frac{hc}{k}$$
 = 1.4387 cm $^{\circ}$ K

4.2.3 JM, Radiative Rate in Band

$$j(T) = \frac{1}{\pi} \int_{\lambda_{\min}}^{\lambda_{\max}} k_{\lambda}(T) R_{\lambda}^{o}(T) d_{\lambda}$$
 (4-15)

where $k_{\lambda}(T)$ = the spectral linear absorption coefficient per molecule from Equation 2-8, multiplied by kT/p to change from k_{λ} per atm.

4.2.4 JT, Radiative Rate Versus Time

$$-\int_{T_{max}}^{T} \frac{dE_{T}}{4\pi j} = t \qquad (4-16)$$

where T is found for value of E_T from results of Subsection 4.2.2 above, and then j is formed from results of Subsection 4.2.3.

4.2.5 Closed Form for Radiance

A least-square fit is made to the resulting data of Subsection 4, 2, 4 of the form

$$j(t) = A(\tau_R - t)$$
 (4-17)

Substituting Equation 4-17 into 4-16 yields

$$\tau_{\mathbf{r}}^2 = \frac{\mathbf{E}_{\mathbf{T}}}{2\pi\mathbf{A}} \tag{4-18}$$

which relates τ_r to T_m , the maximum temperature excited by atmospheric collision.

4.3 PROGRAM SECTION 3. SELECTION OF x OR z_1

These are the initial values of x_1 or z_1 to be used to define lines of sight just grazing plume.

For a \le 45 degrees, we have the following:

$$\theta_{\rm m}$$
 < 90 and 90 - $\theta_{\rm m}$ > α

$$z_{\rm l} = 0 \tag{4-19a}$$

$$\theta_{\rm m}$$
 > 90 and $\theta_{\rm m}$ - 90 < 90 - α

$$z_1 = -r_m/\cos a \qquad (4-19b)$$

Otherwise

$$z_1 = r_{m1} \cos{(\alpha + \theta_m)/\cos{\alpha}}$$
 (4-19c)

For a > 45 degrees, we have the following:

$$\theta_{\mathbf{m}}$$
 < 90 degrees and α > $\theta_{\mathbf{m}}$

$$x_1 = -r_m \cos (a - \theta_m) / \sin a \qquad (4-20a)$$

Otherwise

$$x_1 = -r_m/\sin \alpha \qquad (4-20b)$$

In addition, if there is no gaseous radiation (XM = 0), the plume angular extent is taken equal to the flow angle of the limiting streamline for the smallest particle (33.52 degrees).

4.4 PROGRAM SECTION 4. CALL STAB--INTERCEPT VALUES For $0 \le a < 45 \text{ degrees}$

$$(1 + \tan^2 \alpha) x^2 + (2 z_1 \tan \alpha) x + y_1^2 + z_1^2 - r_m^2 = 0$$
 (3-27, Section 3)

$$(\cot^2 \theta_{\rm m} - \tan^2 \alpha) x^2 - (2 z_1 \tan \alpha) x + y_1^2 \cot^2 \theta_{\rm m} - z_1^2 = 0$$
 (3-28, Section 3)

and for 45 degrees ≤ a ≤ 90 degrees

$$(1 - \cot^2 a) z^2 + (2 x_1 \cot a) z + y_1^2 + x_1^2 - r_m^2 = 0$$
 (3-29, Section 3)

$$(\tan^2 \theta_m - \cot^2 \alpha) z^2 - (2 x_1 \cot \alpha) z - (y_1^2 + x_1^2) = 0$$
 (3-30, Section 3)

4.5 PROGRAM SECTION 5. CALL FUN--RADIANCE CALCULATIONS

$$z = \tan \alpha (x - x_1) + z_1$$
 (3-18, Section 3)

$$r = \sqrt{z^2 + y_1^2 + \left[\cot \alpha (z - z_1) + x_1\right]^2}$$
 (4-21)

$$\theta = \cos^{-1}(z/r) \tag{4-22}$$

4.5.1 Solid Particles

4.5.1.1 Calculate Throat Radius

$$r_t = r_e / \sqrt{\epsilon}$$

where

$$\epsilon = \frac{1}{M_e} \left(\frac{\gamma + 1}{2} \right)^{-\frac{\gamma + 1}{2(\gamma - 1)}} \qquad \left(1 + \frac{\gamma - 1}{2} \quad M_e^2 \right)^{-\frac{\gamma + 1}{2(\gamma - 1)}} \tag{4-23}$$

4.5.1.2 Particle Limiting Velocity

$$V_{(\rho_i, \theta)} = [a_i + b_i \exp(-c_i r/r)] \cos^{n_i} \theta$$
 (4-24)

for $0 \le \theta \le \theta_i$ (limiting particle streamline inclination) where the coefficients are given in Table 4-1.

4.5.1,3 Particle Time of Flight

$$t_{f} = r/V_{(\rho_{i},\theta)} \tag{4-25}$$

4.5.1.4 Concentration

Density ratio =
$$\frac{d_i}{d_0}$$
 = $f(\rho_i, r, \theta) \equiv g_i \left(\frac{r_t}{r}\right)^{e_i} (\cos \theta)^{f_i}$ (4-26)

for $0 \le \theta \le \theta_i$ and the coefficients are given in Table 4-2. Concentration is related to density by

$$N_o = d_o f(\frac{4}{5})\pi \rho_i^3 d'$$
 (XML), $d' = alumina density$
= 4.0045 g/cm³

XML = alumina mass fraction

4.5.1.5 Particle Temperature

If $t_f < \Delta t_{i,s}$ (Equation 4-8),

$$T = T_f$$

If $t_f > \Delta t_{i,s}$,

T = maximum of

a. TCOOL (
$$t_f - \Delta t_{i,s}$$
), Equation 4-3

or b. TMIN, Equations 4-6 or 4-7

4.5.1.6 Evaluate Integrand

$$J_{i} = N_{o} (t' - t_{f}) f(\rho_{i}, r, \theta) \pi \rho_{i}^{2} \epsilon_{i, \Delta \lambda} (T) B_{\Delta \lambda}^{o} (T)$$
(4-27)

(See Equations 4-26, 4-4, and 4-5)

Table 4-1
PARTICLE LIMITING VELOCITY COEFFICIENTS

$ρ_{\mathbf{i}}$ (μm)	$ heta_{ extbf{i}}$	a _i (cm/sec)	b _i (cm/sec)	c _i	n _i
0.690	33,520	30.776×10^4	-4.6692×10^4	0.01874	-0.9048
1.330	28.22	29.701	-3.8094	0.01930	-1.3494
2.000	23.53	28.932	-3.2891	0.01976	-1.5419
2.660	19.76	28.248	-3.0367	0.02012	-1.4886
3.332	16.81	27.520	-2.9355	0.02036	-1.3175
4.000	14.75	26.764	-2.9337	0.02064	-1.0308
4.666	13.56	26.007	-2.9745	0.02092	-0.6425

Table 4-2
CONCENTRATION COEFFICIENTS

$ ho_{f i}^{}(\mu m)$	g _i	e _i	f _i
0.690	0.2201	2.33	2.84
1.330	0.4503	2.24	5.64
2.000	0.5395	2.17	6.75
2.660	1.0308	2.12	7.30
3.332	1.3339	2.08	6.95
4.000	0.8870	2.()	6.30
4.666	0.5396	2.04	-2.05

4.5.2 Excited Gases

4.5.2.1 Time of Flight

$$t_f = r/V_p$$
, Equation 4-9 (4-28)

4.5.2.2 Gas Concentration

$$n_{p} = n_{p,e} \frac{k}{2} \left(\frac{r_{e}}{r}\right)^{2} \left[\cos \left(a\theta\right)\right]^{k-2} \tag{4-29}$$

where

$$k = \gamma (\gamma - 1) M_e^2$$
 (4-30)

$$a = \frac{90 \text{ degrees}}{\theta_e + (\nu_{\text{max}} - \nu_e)}$$
 (4-31)

$$v_{\text{max}} = \left[\sqrt{\frac{Y+1}{Y-1}} - 1 \right]$$
 90 degrees (4-32)

$$v_{\rm e} = \sqrt{\frac{\gamma + 1}{\gamma - 1}} \quad \tan^{-1} \sqrt{\frac{\gamma - 1}{\gamma + 1}} \, (M_{\rm e}^2 - 1) \quad -\tan^{-1} \sqrt{M_{\rm e}^2 - 1} \quad (4-33)$$

and $n_{p,e} = d_e N_o / \overline{M}$

N_o = Avogadro's number

 \overline{M} = mean molecular weight

and remaining parameters are input data

4.5.2.3 Relative Velocity

$$x = \cot \alpha (z - z_1) + x_1$$

$$V_{\infty} = \left[\left(\frac{xV_{p}}{r} + V1 \right)^{2} + \left(\frac{yV_{p}}{r} + V2 \right)^{2} + \left(\frac{zV_{p}}{r} + V3 \right)^{2} \right]^{1/2}$$
 (4-34)

where

 $V1 = V_R \sin \phi_R \cos \psi_R$

 $V2 = V_R \sin \phi_R \sin \phi_R$

 $V3 = V_R \cos \phi_R$

4.5.2.4 Excitation Temperature

T_m = maximum of

A. T of ambient atmosphere (See Equation 4-40)

B.
$$T_{M} = 9.61194 \times 10^{-9} V_{\infty}^{2} + 1.07636 \times 10^{-3} V_{\infty}$$
 (4-35)

(V_∞ in cm/sec)

4.5.2.5 Attenuation Factor

$$A_{t} = S \int_{O}^{r} n_{\infty} (h_{s}) ds \qquad (2-7)$$

where the altitudes, h, are found from

$$h_s = h_o + s \left(\frac{h - h_o}{r}\right)$$

4.5.2.6 Radiative Relaxation Time

$$\tau_{\rm r} = \tau_{\rm r} (T_{\rm m}), \text{ Equation 4-18}$$
 (4-36)

4.5.2.7 Radiance Calculation

$$J_g = V_{\infty} n_{\infty} S \tau_r [1 - \exp(t_f/\tau_r)] n_p j \exp(-A_t) x_n$$
 (4-37)

where for $t_f \ge \tau_r$

$$j = \frac{A\tau}{2}, \quad \text{(See Equation 4-17)}$$

and for $t_f < \tau_r$

$$j = A\left(\tau_r - \frac{t_f}{2}\right) \tag{4-39}$$

 n_{∞} is found from atmosphere table, $n_{\infty}(h)$

S = built-in value of cross section

$$h = h_o + x \sin \phi_u \cos \psi_u + y \sin \phi_u \sin \psi_u + z \cos \phi_u \qquad (4-40)$$

(See Equation 3-40)

4.5.3 Total Radiance

If both particles and gas are included the total integrand is

$$J = J_p + J_g \tag{4-41}$$

4.6 PROGRAM SECTION 6. CALL SIMP--INTEGRATION THROUGH PLUME Simpson's rule

$$\int_{s}^{s+\Delta s} f(x) dx = \frac{\Delta s}{3} \left[f(s) + 4 f\left(s + \frac{\Delta s}{2}\right) + f(s + ds) \right]$$
 (4-42)

Step size is controlled by a geometric progression, so that each step is 10 - larger or smaller than the previous, and minimum step size occurs at z = 0 or x = 0 planes.

4.7 CURVE FITTING ROUTINE, CF

An auxiliary routine was written to obtain a curve fit to the intensity data to facilitate plotting. The curve fit for intensity is of the form

$$I = \sum_{j=1}^{6} a_j \phi_j (x, y)$$

where

$$\phi_{j} = x^{2}, x, y^{2}, y, xy, 1$$

The least square fit condition is (for nine points)

$$\frac{d}{da_k} \sum_{i=1}^{9} \left[I_i - \sum_{j=1}^{6} a_j \phi_j (x_i, y_i) \right]^2 = 0$$

The resulting matrix of equations is

$$(C) (a) = (I)$$

where

and each term in the C and I matrices is summed over all points, e.g.

$$C_{11} = \sum_{i=1}^{9} z_i^4$$

This technique of extrapolation is similar to that used by Dailey (Reference 12). Nine points (3 x 3 square) are selected and the curve fit made. Coordinates of the isointensity lines are then deduced from the curve fit by solving the quadratic equation for y, given I and x. These lines are plotted, and then the next group of nine points are selected. The procedure is repeated until the plume is covered.

Section 5 SAMPLE PROBLEM AND OPERATING INSTRUCTIONS

A list of all input quantities is given in Table 3-1* in Subsection 3.2, with their appropriate units. Additional description of the input parameters especially the Flags, is available in the problem output shown in Appendix B. The input is, in general, standard NAMELIST format (see listing of sample data in Appendix A). Following the NAMELIST input is a single card consisting of a case description (any 80 alphanumeric characters) which is printed at the start of the case. To signify the end of the run (no more cases), a NAMELIST-format input is required setting the flag, END, equal to 1. No case-description card should follow here.

A sample problem follows for an engine typical of a solid-propellant interceptor sustainer rocket. Since the radiating species mole fraction (XM) has been set equal to zero, only radiation from the alumina particles will show up. The specification of look angle equals zero implies a broadside look. The vertical angles equal to 90 degrees means that the plume axis is parallel to the Earth's surface. The values of 90 degrees for the velocity vector imply that the free-stream velocity is perpendicular to the plume axis. The final columns of printed output present x, y coordinates in cm followed by the intensity of w/cm² sr; many zero values of intensity result because the particles cannot expand through the total plume. The output from the sample case follows: (Note: The "error summary" in the output refers to reading in only five values for density versus time instead of maximum allowable value of 50 — no calculation errors result).

All intermediate calculations leading up to calculation of radiance (emissivities, etc.) need not be repeated on successive cases, if not changed. These calculations will be skipped if ISL is entered as -1 (solid propellant) or -2 (gas only)

^{*}Note: If ISL = +1 or -1, the density table (PDE) should be for chamber density and not exit plane density.

Section 6 RECOMMENDATIONS FOR FUTURE WORK

This version of FLAME forms a useful framework for the calculation of plume radiance including a variety of mechanisms. The main accomplishments of this contract were to formulate the integration procedure and geometric-orientation choice to relieve the engineer of tedious extrapolations to attempt to relate flight data to calculations. Because of the large number of options included in FLAME, a comprehensive checkout was not possible during this contract (over 50 runs were made). A streamlining of the output and calculation procedures would be possible through a more extensive exercising of the FLAME options. A variety of values for look angle, altitude, times, orientations engine characteristics, relative velocity, etc. should be used as FLAME input to see if further refinements are needed.

Currently the code calculates only half of the plume in a single run which is sufficient for axisymmetry, such as for particle radiation only. The code should be refined to allow total plume extent to be calculated.

A better contour-plotting routine might be found by limiting the closeness of the isointensity lines, and constraining the curve fit to obtain smoother data. Automatic labeling of the contours would be a nice added feature.

The theoretical models used in FLAME are becoming rapidly antiquated. A molecular radiation model based on non-Boltzmann distributions should be investigated and incorporated.

Additional extensive flow-field revision would be useful based on sophisticated multiphase transient plume codes,

Section 7

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Appendix A LISTING OF FLAME CODE

```
C PROGRAM FLAME, VERSION 1
      PROGRAM P2170(INPUT, OUTPUT, FILMPL, TAPE5=INPUT, TAPE6=OUTPUT,
                   TAPE48=FILMPL+TAPE16)
      DIMENSION PX(3), PY(100), PL(3,100), FC(20), IP2(20)
      DIMENSION POE(50), TID(50),
                                         WLE(8), EPT(20,10), TET(20), TCDOL(
     120,10),TC1(20),BPASS(20),FTL(20,10),TMIN(10),DELT(10),ET(20),IR(20
     2),JM(20),JT(20),TJ(20),XP(8),W(8),RP(10),TCND(20),TC1X(20,10),CASE
     3(10), TIDP(51), PDEP(51), T1(10), T2(10), WP(20), CU(2)
      EQUIVALENCE (TCIX, TIDP), (TCIX(1,4), PDEP), (T1,JT), (T2,JT(11))
      REAL IA, ME, JM, JT, LIP
      INTEGER END
      COMMON/MCG/ JO, IPL, PX, PL, PY, FC, XL, XR, YB, YT, IP, IP2
      COMMON/MSTCF/IA+AL+RD+CONF+RM+X1+Y1+RA
      COMMON/MF/1SL.NP.G.RE.GAMMA, ME.AN.TIG.TID.PDE.NR.XML.DELT.TMIN.TCI
     1,TCOOL, PI, RP, TET, ETL, BPASS, TF, VP, P2, XI2, TR, ET, A, V1, V2, V3, H0, WMOL
     2, XM, XP, W, GL
      DATA XP,W,RA,PI,RD,WP/-.9602898565,-.7966664774,-.5255324099,-.183
     14346425,.1834346425,.5255324099,.7966664774,.9602898565,.101228536
     23,.2223810344,.3137066458,.3626837834,.3626837834,.3137066458,.222
     33810344,.1012285363,90.,3.1415927,57.2957795131,.001,19*1./
      NATA NP, RP/7, 46.604E-5, 40.E-5, 33.315E-5, 26.6E-5, 20.001E-5, 13.3015E
     1-5,6,9007E-5/,B/9,49/
      DATA TR/100.,400.,800.,1200.,1600.,2000.,2400.,2700.,3000.,3300.,
     13500., 3900.,4200.,4500.,5000.,5500.,6000.,7000.,9000.,11000./
      DATA TFT/99..150..200..250..300..350..400..450..500..600..700..
     1800.,1000.,1200.,1400.,1700.,2000.,2500.,3200.,6001./,
     2 TF.DF.CP/2323..4240..3.2/
                  INPUT DATA ****
C **** REGION O.
      NAMELIST /IN/TIG, RE, GAMMA, ME, LIP, TE, AL, HO, VR, PHIU, XIU, TC, PHIR, XIR,
     1 WLMIN, WLMAX, XM, IND, ISL, INT, LINES, NR, TID, PDE, END, IPLOT
      GL = ALOG(10.)
      IPI = 0
    1 READ(5, IN)
      IF (FND.EG.1) STOP
      JP = 1
      KP = 1
      IPL = 0
      00 \ 32 \ I = 1.20
   32 IP2(1) = 0
      IP = 0
      READ(5,200) CASE
      WRITE(6,100) CASE, TIG, RE, GAMMA, ME, LIP, TE, AL, HO, VR, PHIU, XIU, TC, PHIR
     1,XIR,WLMIN,WLMAX,XM,IND,ISL,INT,LINES,NR,IPLOT
  200 FORMAT(10A8)
  100 FORMAT(1H1.3X.67HPROGRAM P2170
                                          FLAME
                                                     ROCKET EXHAUST PLUME
     10PTICAL SIGNATURES ///20x,10A8//11H INPUT DATA//29H TIME AFTER IGN
     2ITION-TIG(SEC)F10.4.11X.25HNNZZLE EXIT RADIUS-RE(CM)F10.4.13X. 5HG
     3AMMAF10.4/ X,19HEXIT MACH NUMBER-MFF10.4,20X,25HNOZZLE LIP ANGLE-
     4LIP(DEG)F10.4.13X.26HEFFECTIVE FARTH TEMP-TE(K)F8.2/ X,18HLOOK AN
     5GLF-AL (DEG) F8.2.23X.22HENGINE ALTITUDE-HO (FT) F9.0.17X.23HROCKET VE
     6LOCITY-VR(FPS)F9.2/ X.27HVERT. POLAR ANGLE-PHIU(DEG)F8.2.14X.30HV
     7ERT. AZIMUTHAL ANGLE-XIU(DEG)F8.2.10X.19HCHAMBER TEMP.-TC(K)F8.2/
     8 X,26HVEL. PULAR ANGLE-PHIR(DEG)F8.2,15X.29HVEL. A7IMUTHAL ANGLE-X
     91R(DEG)F8.2/16H BAND LIMITS AREF6.2.3H TOF6.2.19X.30HRADIATING SPE
     FCILS MASS FRAC-XM F8.5.//
                                                       23H CONTROL FLAGS F
                   IND =15,25H (1 FOR DAY, 0 FOR NIGHT)/8H
                                                                ISL =15.79
     H8/\--WOJJOA
     BH (1,-1 FOR PARTICLES; 0,-2 FOR GAS ONLY; NEGATIVE TO SKIP INITIAL
                                                INT =15.56H (APPROXIMATE
     G CALCULATIONS)
                                          /8H
     CNUMBER OF INTEGRATION STEPS THROUGH LAYER)/8H LINES =15,56H (APPRO
```

STATE OF THE STATE

```
DYIMATE NUMBER OF LINES OF SIGHT IN ONE DIMENSION 1/8H
                                                                NR =15,44H
     E (NUMBER OF ENTRIES IN FIRING HISTORY TABLE)/8H 1PLOT =15,42H (O F
     FOR NO PLOT, 1 FOR ISOINTENSITY PLOTS))
C **** INITIALIZATION OF SD 4020 CAMERA
      IF(IPLOT.NE.O) GO TO 30
   31 \text{ IR} = (NR+2)/3
      K = 1
      00.2.1 = 1.1R
      00 2 J =1.3
      L = I + (J-1)*IR
      TIDP(K) = TID(L)
      PDFP(K) = PDF(I_{L})
    2 K = K + 1
      WRITE (6,101)(TIDP(I),PDEP(I),I =1,NR)
  101 FORMAT(72HO FIRING HISTORY TABLE FOLLOWS, EXIT-PLANE DENSITY (G/CM
     13) VS TIME (SEC) //3(4x,4HTIME,5X,7HDENSITY,5X)/(3(F10.4,E11.3,4X)
C **** REGION 1. INITIAL CALCULATIONS - PARTICLES ****
      CC = (WLMAX - WLMIN)/2.
      DD = (WLMAX + WLMIN)/2.
      IF(ISL.NF.1) GO TO 3
      DO 4 I = 1.20
      T = TEl(I)
C FIND WAVELENGTH WHERE PLANCK FUNCTION IS GREATEST. TRANSFORM INTEGRAL
      WW = 2898 - /T
      C = 3.9*WW
      D = 4.1 = WW
C PERFORM GAUSSIAN INTEGRATION (EQUATIONS 4-1.-2.-4.-5)
      BPASS(I) = 0.
      R1X = 0.
      00.7 K = 1.NP
      EPI(I,K) = 0.
    7 ETL(I,K) = 0.
      00.5 J = 1.8
      WE = C*XP(J) + D
      WWL = CC * XP(J) + DD
      R1 = 3.742E4/WL**5/(EXP(14390./WL/T)-1.)
      R2 = 3.742E4/WWL**5/(EXP(14390./WWL/T)-1.)
      R1X = R1X + W(J) * R1
      BPASS(I) = RPASS(I) + R2*W(J)
C SUM OVER PARTICLES
      DO 5 K = 1.NP
      EM1 = EMIT(KP(K), WL)
      EM2 = EMIT(RP(K)+WWL)
      EPT(I \cdot K) = EPT(I \cdot K) + W(J) *R1 *EM1
    5 ETL(I,K) = ETL(I,K) + W(J)*R2*EM2
      DD 6 K = 1.NP
      EPT(I,K) = FPT(I,K)/R1X
    6 \text{ ETL}(I,K) = \text{ETL}(I,K)/\text{RPASS}(I)
    4 RPASS(1) = RPASS(1)*CC/3.14159
      WRITE(6,103) (RP(I), I=1,NP)
  103 FORMAT(46HO INITIAL CALCULATIONS FOR PARTICLE PROPERTIES//35X.16HT
     10TAL FMISSIVITY//16X,27HPARTICLE RADIUS (CM X 10E4)/13H
     2 .4P10F10.3//)
      00 20 I = 1.20
   20 WRITF(6,104) TET(I),(EPT(I,J),J=1,NP)
  104 FORMAT(1H ,F9.1.3X,10F10.5)
```

```
WRITE(6,107) WEMIN.WEMAX, (RP(I), I=1, NP)
   107 FORMAT(////13HO BAND DATA,(,F5.1,4H TO ,F5.1,1H)//32X,18HEMISSIVIT
      1Y IN RAND//12X,42HPLANCK FUNC. PARTICLE RADIUS (CM X 10E4)/25H
      2 TEMP(K) (W/CM2/SR)
                              ,4P10F9.3//)
       00.22 I = 1.20
    22 WRITE(6,108) TET(1), BPASS(1), (ETL(1,J),J = 1,NP)
   108 FORMAT(1H ,F9.1,F13.4,2X,10F9.5)
C COOLING CURVES FOR SOLID PARTICLES (SECTION 4-1-2-EDS. 4-3 AND 4-8)
       DT = (TF-100.)/19.
       TCOO(1) = TF
       DO 8 K = 1.NP
       TCIX(1,K) = 0.
      T1(K) =]./TF**4*RP(K)
                              /TABLE(TF,TET,EPT(1,K),20,0,1)
    8 DELT(K) = T1(K) *DF/...7016E-11
      T = TF
      00.9 I = 2.20
      T = T - DT
      TCOO(1) = \Gamma
      D0.9 K = 1.NP
      T2(K) = 1./T**4*RP(K)
                             /TABLE(T,TET,EPT(1,K),20,0,1)
      TCIX(1,K) = TCIX(I-1,K)+(T2(K) + T1(K))/2.*DT
    9 T1(K) = T2(K)
      00 10 K = 1.NP
   10^{-1}CIX(20,K) = TCIX(20,1)
C REARRANGE TO CHANGE INDEPENDENT VARIABLE TO TIME.
      DT = TCIX(20.1)/722700.
      TIMF = -DT
      00 \ 11 \ I = 1.20
      TIME = TIME + DT*FL()AT(1**4)
      TCI(I) = TIME/1.7016E-11*CP
      DO 11 K = 1.NP
   11 TCOOL(I_*K) =
                                  TABLE(TIME, TCIX(1,K), TC00,20,0,1)
      TCI(20) = 1.E5
      WRITE(6,105) (RP(I), I=1,NP)
  105 FORMAT(///25X,34HPARTICLE TEMPERATURE HISTORIES (K)//16X,27HPARTIC
     1LE RADIUS (CM X 10F4)/13H TIME(SEC) ,4P10F10.3//)
      00 21 I = 1.20
   21 WRITE(6,106) TCI(I),(TCOOL(I,J),J = 1,NP)
  106 FORMAT(1H ,F10.3,2X,19F10.1)
    3 IF(ISL.LE.0) 60 TO 15
      TO = 400.
      DO 12 K = 1.NP
      TCOOL(20.K) = 70.
C NEWTON-RAPHSON SOLUTION TO ED. 4-7 (AND ED. 4-6)
      C = 5.964E10*TARLF(6000.,TET,EPT(1,K),20.0,1)*FLOAT(IND) +
     1TABLE(TE.TET.EPT(1.K),20.0.1)*TE**4/2.
      00 13 I = 1.100
      TN = (2.*T0 + C /T0**3/TABLE(T0,TET,EPT(1,K),20,0,1))/3.
      IF (AHS(TN-TO).LT..1) GO TO 14
  13 TO = IN
      WRITE (6,102) TN, TD, K
 102 FORMAT(58HOERROR, SOLUTION OF FO.(4-7) DOES NOT CONVERGE. NEW VALU
    1E= G8.2, 12H, OLD VALUE= G8.2,15,18H-TH PARTICLE GROUP)
  14 \text{ TMIN(K)} = TN
  12 CONTINUE
      WRI1E(6,109) (RP(I),I = 1,NP)
 109 FORMAT(////27H PARTICLE RADIUS (CMX10E4) +4P10F10.3)
      WRITE(6,110) (DELT(I),I = 1.NP)
 110 FORMAT(27H SOLIDIFICATION TIME (SEC) .10F10.3)
```

```
WRITE(6.111) (TMIN(1).I = 1.NP)
  111 FORMAT(27H FOUTLIBRIUM TEMPERATURE(K) .10F10.2)
 ***** REGION 2, INITIAL CALCULATIONS - GAS ****
 IF RAND LIMITS UNCHANGED FROM PREVIOUS RUN. SKIP BAND-FNERGY CALC.
   15 IF(ISL.L).0) GO 10 16
C DETERMINE ET (FO. 4-11)
      O1 = 0.5E4/R/WEMAX
      Q2 = 0.5E4/R/WEMIN
      00.18 I = 1.20
      T = TR(I)
      SIG = 1.4387#R/T
      YMIN = 01#01#SIG
      112*C0# 20 = XMAX = 02*02*51f
                             9.928E-4*B/SIG*(01*EXP(-XMIN) - 02*EXP(-XMAX
      ET(1) =
     1) + .88623/50RT(SIG)*(ERF(02*SQRT(SIG)) - ERF(01*SQRT(SIG))))
C DETERMINE BAND RADIANCE (ED. 4-15)
      JM(1) = 0.
      DO 17 J = 1.8
      WWL = CC*XP(J) + DD
   17 JM(T) = JM(T) + W(J) \pm \Delta BCO(WWL T)/WWL \pm 5/(EXP(14390 JWWL/T)-1.)
   18 \text{ JM}(1) = \text{JM}(1) *CC*3.742E4/P1
      WRJ1F(6.112) (TR(I), ET(I), JM(I), I = 1.20)
  112 FORMAT(////43HOROTATIONAL RADIATION FROM WATER MOLECULES./97H
     IAVAILABLE BAND ENERGY = ET (106-20 W SEC/MOLEC), BAND RADIATIVE RA
     2TE = J (10E-20 W/SR MOLEC) //24X,4HT(K),8X,2HET,12X,1HJ//(22X,F7.1
     3,2×,2613.31)
C CALCULATION OF RADIATIVE COOLDOWN (EQ. 4-16) (J VS TIME)
      ETX = ET(1)
      T0 = 0.
      DEI = [T(20)/190.
      JT(1) = JM(1)
      TJ(1) = 0.
      00.19 I = 2.20
      ETX = ETX +DET
      JT(I) = TABLE(ETX-ET,JM,20,0,1)
      TN = .0/9578/JT(I)

TJ(I) = TJ(I-1) + (TO + TN)/2.*DET
   19 TO = TN
C FIT DATA WITH FUNCTION LINEAR IN TIME (EQ. 4-17)
      CALL LSOPOL(TJ,JT,WP,TIDP,20,ETX,2,CU)
      A = CU(2)
      00 23 : = 1.20
      TCOO(1) = CU(1) + CU(2)*IJ(1)
      TIDP(I) = TABLE(JT(I), JM, ET, 20, 0, 1)
   23 PDEP(I) = TA9LE(JT(I), JM, TR, 20, 0, -1)
      WRITE(6,113) (T.I'I),JT(I),TCOO(I),TIDP(I),PDEP(I),I = 1,20)
                          RADIATIVE DECAY IN BAND //9X, 9HTIME(SEC),6X,1H
  113 FORMAT(////28H
     1J.8X,6HJ(FIT),7X,2HET,8X,4HT(K)//(6X,F11.5,2X,4E11.3))
 SOLUTION FOR LIMITING VELOCITY (EQS. 4-9 AND 4-10)
   16 VP = 2.*SORT(TC/WMOL/(2.-2./GAMMA)*8.3144E7)
C PREPARATION FOR PLUME INTEGRATION
      X1 = 0.
C CALCULATION OF MAXIMUM PLUME CONE HALF-ANGLE AND MAXIMUM PLUME RADIUS
      GG = SORT((GAMMA+1.)/(GAMMA-1.))
```

```
GM = SORT(ME**2 - 1.)
      CONE = LIP + RA*(GG-1) -
                                        RD*(GG*ATAN(GM/GG) +ATAN(GM))
      RM = VP*TIG
      DELS = (RM + RM*COS(AL/RD))**05/(1.1**(INT/2) ~ 1.)
      IA = 0.
          = PHIR/RD
      Pl
      P2
          = PHIU/RD
      XII = XIR/RD
      XI2 = XIU/RD
      IF(\Delta L \cdot GT \cdot 45 \cdot) IA = 1.
      DELX = DELS * COS((RA * IA - AL)/RD)
      CELY = (RM + RM*SIN(AL/RD))*.1/(1.1**LINES - 1.)
      CD = COS((RA*IA-AL)/RD)
      DELXX = DELY/CD
      DEL = DELY
      J0 = 10000
C SELECTION OF ISOINTERSITY LEVELS TO BE PLOTTED
      FC(1) = 1 \cdot E - 11
      X10 = SORT(10.)
      D0 29 I = 1.19
   29 FC(1+1) = FC(1) * X 10
 ***** REGION 3. SELECTION OF INITIAL VALUE OF X1 *****
C
      VV = VR*30.4801
      V1 = VV*SIN(P1)*COS(XI1)
      V2 = VV \pm SIN(P1) \pm SIN(XI1)
      V3 = VV*COS(P1)
      INOUT = 2
      IF(14.EC.1.) GO TO 36
      1F(XM.EQ.O.) GO TO 27
      IF(CONE.GT.90.) GO TO 25
      1F(90. - CONE.GF.AL) GO TO 27
   34 \times 1 = RM \times COS((CONE + AL)/RD)/COS(AL/RD)
      GO TO 27
   25 IF(90. - CONE.GE.AL - 90.) GO TO 34
      X1 = -RM/COS(AL/RD)
      GO TO 27
   36 IF(XM.NE.O.) GO TO 33
      X1 = -RM = COS((AL - 33.52)/RD)/SIN(AL/RD)
      GU TU 27
   33 IF(CONE.LT.90..AND.CONE.LT.AL) GO TO 35
      X1 = -RM/SIN(AL/PD)
      GO TO 27
   35 XI = -RM * COS((AL - CONE)/RD)/SIN(AL/RD)
   27 XPLOT = 0.
      PX(1) = XPLOT
      DELX1 = (DELXX + ABS(X1)/10.)/1.1
      R = 1./1.1
C PLOT LIMITS FOR GRID
      XL = 0.
      XR = RM - X1 + COS((RA + IA - AL)/RD)
      IF(IA.EO.1..AND.XM.EQ.O.) XR = 2.*RM*SIN(33.52/RD)
     1 *SIN(AL/RD)
      YB = 0.
      YT = XR
   **** REGION 4. CALCULATE INTEGRATION BOUNDARIES ****
```

```
24 CALL STAR (PMIN, PMAX, INOUT)
      DELY = DELY*1.1
      IF(INOUT.FO.O) GO TO 26
      Y1 = 0.
      DELY = DEL
      JO = MINO(JO,JP)
      KP = KP + 1
      IF(KP.EQ.4.AND.IPLOT.NE.O) CALL CONTUR(KP)
      JP = 0
      IF (X1.GE.0.) R = 1.7
      DEL XI = DEL XI *R
      X1 = X1 + DELX1
      XPLOT = XPLOT + DELX1*CD
C STURE VALUES (PX.PL.PY) FOR PLOTTING
     PX(KP) = XPLOT
      CALL STAB (PMIN, PMAX, INOUT)
      IF(INOUT.EQ.O) GO TO 26
      GU TO 1
C **** REGION 5. INITIAL EVALUATION OF INTEGRAND ****
C
  26 CALL FUN(SOLD, PMIN)
C **** REGION 6, INTEGRATE THROUGH PLUME *****
      CALL SIMP(SOLD, SINT, PMAX, PMIN, DELX)
      SINT = SINT *.6666667
      WRITE(6,114) XPLOT, Y1, SINT
  114 FORMAT(4X,2F10.2,E11.3)
      JP = JP + I
      PL(KP, JP) = SINT
      PY(JP) = Y1
      Y1 = Y1 + 0ELY
      GO TO 24
   30 IF(IPT.EQ.O) CALL CAMRAV(IPLOT)
      IPT = 1
      GO 10 31
      END
```

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SUBROUTINE SIMP(SOLD, SINT, PMAX, PMIN, DEL)
  SINT = SOLD
  DELX = (ABS(PMIN) /10. + DEL)/1.1
  A = 1./1.1
  X = PMIN
1 SUM = 0.
  c = 4.
  DELX = DELX#A
4 X = X + DFLX
  IF(X.GT.PMAX) GO TO 2
  CALL FUN(SS,X)
  SS = SS*DELX
SUM = SUM + C*SS
IF(C.E0.4.) GO TO 3
  IF(X.GT.O.) GO TO 6
5 SOLD = SS
  SINT = SINT + SUM
  GO TO 1
3 C = 2.
  GD TO 4
2 SINT = SINT - SOLD
  RETURN
6 A = 1.1
GO TO 5
  END
```

```
SUBROUTINE STAB (PMIN, PMAX, INOUT)
      DIMENSION X(2)
      COMMON/STC/X+
      COMMON/MSTCF/1A +AL +RD +CONE +RM+X1 +Y1 +RA
      REAL IA
      IF(INOUT.NE.2) GO TO 1
      \Delta O = \Delta BS (TAN((RA # IA - AL)/RD))
      \Delta 1 = \Delta 0 = \Delta 0
      \Delta 2 = (TAN((RA*(IA-1.)-CONE)/RD))**2
    1 \text{ INOUT = 0}
      81 = X1*A0/(1. + A1)
      B = Y1 **7
      C1 = B + X1 **2
                          --RM≄RM
      IF(I\Lambda \cdot EQ \cdot O \cdot)B = -B * \Lambda 2
      C2 = X1 * X1 + B
      RAD = B1 * B1 - C1/(1.4 A1)
      JI = 1
C CHECK FOR LINE-OF-SIGHT BEYOND R=RP.
      IF(RAD.LT.O.) GO TO 5
      RAD = SQRT(RAD)
      XX = -R1 + RAD
      CALL CHECK(II)
      IF(II.6T.2) GO TO 4
      XX = -81 - RAD
      CALL CHECK(II)
      IF(II.GT.2) GO TO 4
    3 RD = A1 - A2
      B2 = X1 * AO/BD
C CHECK FOR LINE-OF-SIGHT BEING PARALLEL TO PLUME BOUNDARY.
      IF(ARS(BD).LT.1.E-8) GO TO 2
      RAD = B2*B2 - C2/BD
      IF (RAD.LT.O.) GO TO 5
       RAD = SQRT(RAD)
       XX = -82 + RAD
      CALL CHECK(II)
       IF(11.GT.2) GO TO 4
       XX = -82 - RA0
    7 CALL CHECK(II)
       IF(II.GT.2) GO TO 4
     5 INOUT = 1
       RETURN
    2 IF(AL.EQ.O..(OR.AL.EQ.RA) GO TO 5
       XX = C2/B2
       GO TO 7
     4 IF (X(1).GT. X(2)) GO TO 6
       PMIN = X(1)
       PM\Delta X = X(2)
       RETURN
     6 \text{ PMIN} = X(2)
       PMAX = X(1)
       RETURN
       END
```

```
SURROUTINE CHECK(I)
      DIMENSION X(S)
      COMMON/STC/X.
      COMMON/MSTCF/IA+AL+RD+CONE+RM+X1+Y1+RA
C CHECK TO SEE IF XX REPRESENTS X OR Z.
      IF(IA.EQ.O.) 60 TO 1
      77 = XX
              SORT(Y1*Y1 + (ZZ*TAN((RA-AL)/RD) + X1)**2 +ZZ*ZZ)
    2 IF (RR.EQ.O.) RETURN
      THEIR = RD*ACOS(77/RR)
C CHECK TO SEE IF INTERSECTION IS ON CONICAL PLUME BOUNDARY.
      IF (ARS) (THETA-CONE)/CONE).LT..OO1) GO TO 3
C INTERSECTION IS ON SPHERICAL CAP OF PLUME.
      IF (THETA.GT.CONE) RETURN
C DO THE FOLLOWING IF THIS POINT IS INDEED ONE OF THE INTEGRATION LIMITS
    4 X(I) = XX
      I = 1+1
      RETURN
    1 \ ZZ = XX*TAN(AL/RD) + X1
      RR = SORT(Y1*Y1 + XX*XX + ZZ*ZZ)
      GO TO 2
C INTERSECTION IS ON CONICAL BOUNDARY
    3 IF(RR/RM.LT.1.001) GO TO 4
      RETURN
      END
```

```
SUBROUTINE FUN(0,P)
      DIMENSION PDF (50), TID (50), TET (20), TC (00L (20,10), TC I (20), BPASS (20),
     1FTL(20,10),TMIN(10),DELT(10),FT(20),TR(20),RP(10),N(10),AB(10),B(1
     20),C(10),TH(10),D(10),E(10),F(10),H1(20),TMT(20),CAT(20),XP(8)
     3.W(R)
      REAL ME, TA, J, JG, N
      CUMMON/ME/ISC: NP.R.PE.GAMMA.ME.AN.TIG.TID.PDE.NR.XML.DELT.TMIN.TCI
     1,TCOUL, 11.RP., TET., BPASS, TF., VP., PHIU, XIU, TR, ET, A, V1, V2, V3, H0, WMOL
     2. YM, XP, W.GL
      COMMUNIMSTOF/IA + AL + RD + CONF + RM + X1 + Y1 + RA
      DATA WMOL.XML.DP/17. . .36 .4.0045/.S.CF/.7E-15.3.531445E-5/
C CURVE FIT COFFFICIENT V AND RHO CORRELATIONS FOR TWO-PHASE FLOW
      DATA AB/2.6007F5.2.6764E5,2.7520F5,2.8248E5,2.8932E5,2.9701F5.3.07
     176F5/,8/-2.9745F4,-2.9337E4,-2.9355F4,-3.0367E4,-3.2891E4,-3.8094E
     74,-4.6692F4/.C/.02092,.02064,.02036,.02012,.01976,.01930,.01874/,N
     3/-.6425,-1.0308,-1.3175,-1.4886,-1.5419,-1.3494,-.9048/,D/.5396,.8
     487.1.3339.1.0308..5395..4053..2201/.E/2.04.2.06.2.08.2.12.2.17.2.2
     54.2.33/.F/-2.05.6.30.6.95.7.30.6.75.5.65.2.84/.TH/13.56.14.75.16.8
     61,19.76,23.53,28.22,33.52/
C H(FT), T, AND LOG(N/FT3) FROM U.S. STANDARD ATMOSPHERE (1962)
      DATA HT+TMT+CAT/3.F5+3.5F5+4.E5+4.5E5+5.F5+5.5F5+6.F5+6.5E5+7.E5+8
     1.65,9.65,1.66,1.166,1.366,1.566,1.766,1.966,2.166,2.366,3.066,184.
     295,241,74,385,34,663,20,924,16,1084,09,1170,47,1230,3,1275,21,1346
     3.67,1396.14,1435.68,1455.53,1485.69,1489.35,1496.29,1500.57,1506.3
     43.2*1507.58,18.10,16.98,16.05,15.41,15.01,14.75,14.55,14.36,14.20,
     513.90,13.64,13.40,13.17,12.76,12.41,12.08,11.76,11.49,11.20,10.3/
      X = P
      JG = 0.
      S1 = 0.
      G2 = G\Delta MM\Delta - 1.
      Z = X1 + ABS(TAN((RA*IA-AL)/RD))*X
      IF(14.EQ.1.) GD TO 6
    7 R = SQRT(X*X + Y1*Y1 + Z*Z)
      \Delta N = \Delta COS(Z/R) *RD
      IF(IABS(ISL).NE.1) GO TO 1
      G1 = (G\Delta MM\Delta + 1.)/2.
C CALCULATE THROAT RADIUS FROM EXIT RADIUS AND EXPANSION RATIO
      RT = RE/SORT(G1**(-G1/G2)*(1.+G2*ME*ME/2.)**G1/G2/ME)
      CT= COS(AN/RD)
      00 \ 2 \ I = 1.NP
C CALCULATE PARTICLE TIME OF FLIGHT. FO. 4-25
      IF(TH(1).LT.AN) GO TO 2
C ARE WE BEYOND THE LIMITING STREAMLINE FOR THIS SIZE PARTICLE
      T1 =
                      R/CT**N(I)/(AB(I)+B(I)*EXP(-C(I)*R/RT))
      IF(T1.GT.TIG) GO TO 2
      PDO = TARLF(TIG-TI+TID+PDE+NR+0+1)*XML
C PARTICLE CONCENTRATION
      PD = PD0*D(1)*(RT/R)**E(1)*CT**F(1)/4.18879/DP/RP(1)**3
C CHECK FOR INCOMPLETE SOLIDIFICATION OF PARTICLE
      IF(TI-LT-DELT(I)) GO TO 4
      T = \Delta MAXI(TMIN(I),TARLE(TI -DELT(I),TCI,TCOOL(1,I),20,0,1))
                                  .20.0.1)
    3 BO = TABLE(T.TFT, BPASS
C PARTICLE RADIANCE ACCUMULATION, EQ. 4-27
      S1 = S1 + PD*PI*RP(1)**2 * FARLE(T, TET, ETL(1,1),20,0,-1)*B0
    2 CONTINUE
      GU TO 1
     T = TF
      GO TO 3
C GAS-PHASE CALCULATIONS.
```

```
1 TI= R/VP
       IF(TI.GT.TIG.OR.AN.GE.CONE) GO TO 5
       GDO = TABLE(TIG-TI, TID, PDE, NR, 0, 1) *6,0254E23/WMOL
       XK = CAMMA*G2*MF*ME
C GAS DEMSITY, FO 4-29 (MODIFIED FOR TWO-PHASE FLOW)
      GD = GDO*xK/?*(RE/R, **2*XM*(COS(RA/CONE*AN/RD))**(XK-2*)
       IF(GD .LT.1.)GO TO 5
      VX = X \neq VP/R
       VZ = Z*VP/R
C RELATIVE VELUCITY, EQ 3-16
      VREL = SORT((VX+V1)**2 + (VZ+V3)**2 + (Y1*VP/R +V2)**2)
C FIND ALTITUDE IN PLUME EQ. 4-40
      H = HO + (SIN(PHIU))*(X*COS(XIU)) + Y1*SIN(XIU)) + Z*COS(PHIU))/30.48
      H = \Delta M \Delta X (H_{*} 3.E5)
      H = \Delta MINI(H+3.E6)
C FIND EXCITATION TEMPERATURE, EQ. 4-35
      TA = TABLE(H,HT,TMT,20,0,1)
       TM = AMAX1(TA, VREL*(9.6)194E-9*VREL + 1.07636E-3))
C ATMOSPHERIC CONCENTRATION
      CA = (EXP(TABLE(H, HT, CAT, 20, 0, -1) *GL)) *CF
C CALCULATION OF ATTENUATION FACTOR. EQ. 2-7
      HD = (H - HO)/2
       . O = TA
      00 \ 8 \ I = 1.8
      HS = HO + HD*(XP(I) + 1.)
    8 \Delta T = \Delta T + W(1) *EXP(TABLE(HS, HT, CAT, 20, 0, 1) *GL)
      \Delta T = \Delta T * R / 2 * "CF * S
C RADIATIVE RELAXATION TIME, EQ. 4-18 AND 4-36
       EM = TABLE(TM,TR,ET,20,0,1)
      TAU = SQRT(EM/2./PI/A)
C COMPUTATION OF EMISSION, EQ. 4-38 AND 4-39
      TX = 1. - TI/TAU
       J = A * TAU/?
       IF(TX \cdot GT \cdot O \cdot) J = J + J*TX
C RADIANCE CALCULATION. FQ. 4-37
      JG = VREL*CA*S*TAU*(1.-EXP(-TI/TAU))*GD*J*1.E-20*EXP(-AT)
    5 \Omega = S1 + JG
      RETURN
    6 TX = Z
      Z = X
       X = TX
      GO TO 7
       END
```

```
FUNCTION ARCO(WL.TE)

C COMPUTES ABSORPTION COEFFICIENT PER MOLECULE OF WATER IN THE

ROTATIONAL SPECTRUM, SCALED BY 10E20

IF(f1.F0.TF) GO TO 1

T1 = TF

T = SORT(T1)

WO = 10.6*T

YO = 279.96/(T + 50.031) - 3.208

A = 1.0303F-4 - 105.582E-4/(T-2.896)

1 W = 10000./WL

ABCO = EXP(2.302585*(YO + A*(W - WO)**2/SORT(W)))*1.3803E-2*TE

RETURN

END
```

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FUNCTION EMIT(R,W)

C COMPUTES THE EMISSIVITY OF ALUMINA PARTICLES OF RADIUS R (CM) AT A C, WAVELENGTH W (MICRONS) BASED ON CORRELATION OF MIE CALCULATIONS.

IF (A.EO.O.) GO TO 1

IF (R.EO.RO) GO TO 2

1 A = (11.32 - 2220.*R) *R + .00689

RO = R

2 EMIY = A*(W + 2.5)

RETURN
END

```
SUBROUTINE CONTUR(KP)
      DIMENSION TT(9', xP(21), YP(21), PX(3), PY(100), PL(3,100), FC(20), CA(6)
     1,182 (20)
      DATA TT/9*10H
      COMMUN/MCG/ JU, TPL, PX, PL, PY, FC, XL, XR, YB, YT, IP, IP2
      J0 = (1J0 + 1)/2)*2 - 1
      IF (IPL.FQ.O) CALL GRID
      J1 = J0 - 2
      FD = 0.
      DX = (PX(3) - PX(1))/20.
      00 1 1, = 1 + J1 + 2
C CURVE FIT TO 3 X 3 MATRIX OF PLUME POINTS
      CALL CF(PX.PY(L),PL(),L),CA)
C HERF TO STATEMENT 7 -- FINDS WHAT ISOINTENSITY LINES ARE IN MATRIX
      AIG = 0.
      SMALL = 1.F20
      12 = 1.+2
      00 \ 2 \ 1 = 1.3
      DO 2 J = L.L2
      A = P((I,J)
      IF(A.FG.O.) GO TO 2
      BIG = AMAXI(RIG \cdot A)
      SMALI = AMINI(SMALL+A)
    2 CONTINUE
      YS = PY(L)
      YL = PY(L2)
      00.9 I = 1.20
      IF (FC(I).LT.SMALL.OR.FC(I).GT.BIG) GO TO 9
      IF (FD.GT.FC(1)) GO TO 7
      FD = FC(1)
      I0 = I
      N = \Gamma
C HERF TO STATEMENT 3 -- SOLVES FOR COORDINATES OF ISOINTENSITY LINES
    7 K = 0
      X = PX(1)
      100 3 J = 1,20
      B = C\Delta(4) + C\Delta(5) * X
      A = CA(3)
      C = CA(1)*X**? + CA(2)*X - FC(1) + CA(6)
      D = B = +2 - 4. +4. +C
      IF (D.LT.O.) GO TO 4
      D = SORT(D)
      Y = (-B + D)/2 \cdot /\Lambda
      IF(Y.GT.YS.AND.Y.LT.YL) GD TO 5
      Y = (-B - D)/2 \cdot /A
      IF(Y.GT.YS.AND.Y.LT.YL) GO TO 5
     x = x + Dx
    3 CONTINUE
      IF (K.LE.1) GO TO 9
      CALL AICRT3(0,0,XP,YP,K,1,1,1,58,TT,TT,TT,2,1,16.,16.,2,XL,XR,2,
     1YB,YT)
    9 CONTINUE
    1 CONTINUE
      00.6 M = 1.1P
    6 IF(IP2(M).FQ.ID) GO TO 10
      IP = IP + 1
      IP2(IP) = I0
      WRITE (6,100) FD ,PX(1),PY(N)
 100 FORMAT(22H ISOINTENSITY LINE FOR E4.2,18HW/CM2/SR IS NEAR (E9.2,1H
```

SURROUTINE GRID

C THIS ROUTINE PREPARES THE GRAPH LIMITS FOR PLOTTING ISDINTENSITY LINES DIMENSION TITLE(8), XT(9), YT(9), PX(3), PY(100), PL(3,100), FC(20) 1, 1P2(20)

COMMON/MCG/ JO, IPL, PX, PL, PY, FC, XL, XR, YR, YP, IP, IPP

DATA TITLE/3*IH , 6HPLUME , 6HCONTOU, 2HRS/, XT/4*IH , 6HX (CM)/, 1 YT/4*IH , 6HY (CM)/

CA! L AICRT3(0,0,X,Y,2,3,1,1,42,TITLE,XT,YI,1,1,16,,16,,2,XL,XR,2,1YB,Y2)

IPL = 1

RETURN
END

```
SURROUTINE CF(X,Y,F,C)
C THIS ROUTINE PERFURMS A LEAST-SQUARE CURVE FIT TO 9 (X,Y) POINTS OF
C FUNCTION F, WHERE
         F(F(T)) = C(1)X + 2 + C(2)X + C(3)Y + 2 + C(4)Y + C(5)XY + C(6)
        DIMENSION X(3),Y(3),F(3,3),C(6),A(6,6),RT(18),IT(18)
        00 \ 1 \ I = 1.6
        C(1) = 0.
        001 J = 1.6
     1 A(I+J) = 0.
        100 ? 1 = 1.3
        X2 = X(1) **2
        \Delta(1,1) = \Delta(1,1) + X2*X2
        \Delta(1,2) = \Delta(1,2) + X2*X(1)
        \Delta(2,6) = \Delta(2,6) + X(1)
     2 \Lambda(2,7) = \Lambda(2,2) + X2
        00 \ 3 \ I = 1.3
        Y2 = Y(1) **2
        \Delta(3,3) = \Delta(3,3) + Y2*Y2
        \Delta(3,4) = \Delta(3,4) + Y2*Y(1)
        \Delta(4,6) = \Delta(4,6) + Y(1)
     3 \Lambda(4,4) = \Lambda(4,4) + Y2
        \Lambda(1,1) = 3.*\Lambda(1,1)
        \Lambda(1,2) = 3.*\Lambda(1,2)
        \Delta(2,2) = 3.*\Delta(2,2)
        \Delta(3,3) = 3.*\Delta(3,3)
        A(3,4) = 3.*A(3,4)
        \Delta(4,4) = 3. \pm \Delta(4,4)
        A(4,6) = 3.*A(4,6)
        \Delta(2,6) = 3.*\Delta(2,6)
        A(2,1) = A(1,2)
        \Delta(4,3) = \Delta(3,4)
        00 \ 4 \ I = 1.3
        DO 4 J = 1.3
        x2 = x(1) * *2
        Y2 = Y(J) **2
        XY = X(I) # Y(J)
        \Delta(1.3) = \Delta(1.3) + 22*Y2
        \Delta(1,4) = \Delta(1,4) + X2*Y(J)
        \Delta(1,5) = \Delta(1,5) + X2*XY
        \Delta(2,3) = \Delta(2,3) + XY + Y(J)
        \Delta(2,4) = \Delta(2,4) + XY
        \Delta(3,5) = \Delta(3,5) + XY*Y2
        C(') = C(1) + F(I+J) * X2
        C(2) = C(2) + F(I_*J)*X(I)
        C(3) = C(3) + F(1,J)*Y2
        C(4) = C(4) + F(1,J)*Y(J)
        C(6) = C(6) + F(I_{3}J)
      4 C(5) = C(5) + F(I \cdot J) * XY
        \Lambda(3,1) = \Lambda(1,3)
        \Delta(3,2) = \Delta(2,3)
        \Delta(4,1) = \Delta(1,4)
        \Delta(4,2) = \Delta(2,4)
        \Delta(4,5) = \Delta(2,3)
        \Delta(5,4) = \Delta(2,3)
        \Delta(5,1) = \Delta(1,5)
        \Delta(5,2) = \Delta(1,4)
        A(2,5) = A(1,4)
        \Delta(5,3) = \Delta(3,5)
```

A(6,6) = 9.

```
A(1.6) = A(2.2)

A(6.1) = A(2.2)

A(6.2) = A(2.6)

A(3.6) = A(4.4)

A(6.3) = A(4.4)

A(6.4) = A(4.6)

A(5.5) = A(1.3)

A(5.6) = A(2.4)

A(6.5) = A(2.4)

CALL SID(A.6.6.6.C.1.1.S.IR.RT.IT.SC)

IF(S.LT.3.) WRITE(6.100) X(1).Y(1)

100 FORMAT (38H CURVETIT COFFFICIENTS INACCURATE AT (E9.2.2H, E9.2.1H)

11

RETURN

END
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THE THE PARTY OF T

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FUNCTION ERF(ETA)

SEE HANDBOOK OF MATHEMATICAL FUNCTIONS WITH FORMULAS, GRAPHS AND MATHEMATICAL TABLES. NBS, APPLIED MATHEMATICAL SERIES 55, PAGE 299, SECTION 7.1.26 (1964)

DATA P/ 0.3275911/

DATA A1,A2,A3,A4,A5 / 0.254829592,-0.284496736,1.421413741, 1-1.453152027,1.06140543 /

T = 1.0/(1. + P*ABS(ETA))

S = T*(A1 + T*(A2 + T*(A3 + T*(A4 + A5*T))))

ERF = 1. - S*FXP(-ETA**2)

IF (FTA.LT. 0.0) ERF = -ERF

RETURN

END
```

```
0010
    FUNCTION TAPLE(XX,X,EX,MO,L,IEX)
                                                                                 0020
    DIMENSION X(MO), EX(MO)
                                                                                  0030
    M = MO
                                                                                 0040
    IF (XX-λ(N)) 1,1,3
                                                                                  0059
    IF (XX-X(1)) 2,6,6
                                                                                 0060
   TABLE = FX(1)
                                                                                  0070
    GO TU 4
                                                                                 0087
  3 TARLE=FX(M)
                                                                                  0090
  4 WRITF(6,5)
                                                                                 0100
  5 FORMAT(1HO,10Y,3MH*** ERROR, RANGE OF TABLE EXCEEDED *** )
    WRITE (6,101) XX,X(1),X(M),MO
101 FORMAT ( 3E13.6.110)
                                                                                  0116
    IF (1EX.50.0) GO TO 100
                                                                                  0120
    STUP
  6 IF (IEX.11.0) GO TO 14
    M1 = 1
                                                                                  0146
    M3 = M
                                                                                  0150
    M2=M/2
                                                                                  0166
  7 1F (X(M2)-XX) 9,8,8
                                                                                  0170
  8 M3=ii2
                                                                                  0180
    GO TO 10
  9 M1=M2
                                                                                  0196
                                                                                  0200
 10 J = (M1 + M3)/2
                                                                                  0210
    IF (J.EO.M2) GO TO 11
                                                                                  0220
    M2 = J
                                                                                  0230
    60 TO 7
                                                                                  0240
 11 IF (X(J)-XX) 12,12,13
                                                                                  0250
 12 M3=J+1
                                                                                  0260
    M1=J
                                                                                  0270
    60 TO 14
                                                                                  0280
 13 M3=J
                                                                                  0290
    M1 = J - 1
 14 IF (L.NE.O) GO TO 15
                                                                                  0300
                                                                                  0310
    TABLE=FX(M1)+(XX-X(M1))*(FX(M3)-FX(M1))/(X(M3)-X(M1))
                                                                                  0320
    GO TO 100
                                                                                  0330
 15 IF (M1.GT.1) GO TO 16
                                                                                  0340
    11=1
                                                                                  0350
    12=2
                                                                                  0360
    13 = 3
                                                                                  0370
    60 10 17
 16 IF (M.GT.M3) GO TO 18
                                                                                  0380,
                                                                                  0390
    13=M
                                                                                  0400
    12 = N-1
                                                                                  0410
    I1 = M - 2
                                                                                  0420
 17 DF32=FX(I3)-FX(I2)
                                                                                  0430
    DF21=FX(I2)-FX(I1)
    DENOM=(X(13)-X(12))*(X(13)-X(11))*(X(12)-X(11))
                                                                                  0440
    Δ=((X(I2)-X(I1))*DF32-(X(I3)-X(I2))*DF21)/DENOM
                                                                                  0450
    B=((X(I2)**7-4(I1)**2)*DF32-(X(I3)**2-X(I2)**2)*DF21)/DENOM
                                                                                  0460
    TABLE=FX(13)-A*(X(13)**2-XX**2)+B*(X(13)-XX)
                                                                                  0470
                                                                                  0480
    GO TO 100
 18 [4=M3+1
                                                                                  0490
                                                                                  0500
     I3=M3
    I2=M1
                                                                                  0510
                                                                                  0520
     I1 = M1 - 1
    DF43=FX(14)-FX(13)
                                                                                  0530
                                                                                  0540
    DF32 = FX(13) - FX(12)
                                                                                  0550
    0F21=FX(I2)-FX(I1)
    \Delta 1 = (X(13) - X(12)) * (X(13) - X(11)) * (X(12) - X(11)) * DF43
                                                                                  0560
```

$\Delta 2 = (X(14) - X(13)) * (X(12) - X(11)) * (X(14) + X(13) - X(12) - X(11)) * DF32$	0570
$\Delta 3 = (X(14) - X(13)) * (X(14) - X(12)) * (X(13) - X(12)) * DF21$	0580
DFNGM=(X(14)-X(13))*(X(14)-X(12))*(X(14)-X(11))*	0590
(x(13)-x(12))*(x(13)-x(11))*(x(12)-x(11))	0600
$\Lambda = (\Lambda 1 - \Lambda 2 + \Lambda 3) / DEN(W)$	0610
$B_1 = (X(13) - X(12)) \approx DE43 - (X(14) - X(13)) *DE32$	0620
B2=(X(14)-X(13))*(X(13)-X(12))*(X(14)-X(12))	0630
$R = R_1 / (2 - h^* (X(14) + X(13) + X(12))$	0640
$C = DF43/(X(14) - Y(13)) - \Delta \div (X(14) * \div 2 + X(13) * \div 2 + X(14) \div X(13))$	0650
1 = R = (X(14) + Y(14))	0660
D=FX(]4)-(A:X(]4):.2+8:X(]4)+C)*X(]4)	0670
TARL 6≈A: УУЖАЗ+R 4×ХЖЖ2 +C*XX+D	0680
100 RETURN	0690
END	0700

\$IN TIG= 1.,RF = 5.7.GAMMA = 1.33, ME = 4.,LIP =15.,TE = 250.,AL = 0., HO= 5.E5, VR = 2.E4,PHIU = 90.,XIU = 90.,PHIR = 90.,XIR = 90.,TC=1255., WLMIN=8..WLMAX=14., IND =0, ISL =1, INT = 20, LINES = 30, NR = 5, TID= -.1,0.,2.,2.1,10.,PDE = 0.,2*1.1E=3,2*0., IPLOT = 1.XM = 0.0 \$ BROADSIDE LOOK, PARTICLE ONLY, 600 LB THRUST, AT NIGH; \$IN END = 1\$

Appendix B OUTPUT FOR SAMPLE PROBLEM

	12 12 12 12 12 12 12 12	: = 7	PA LOADER	FTA LOADER 07471 FT	FWA TABLES	066623	LAGELEDCOPMON	2		
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01736 01736 017363 0173	01780	. 4	027	012216						
017366 017362 017363 017364 017364 02107 0	017366 017362 017363 01	*	111	017324				•	1	
017622 017741 017741 021006 021100 022100 022100 022007 0230007 023007 023007 023007 023007 023007 023007 023007 0230007 023007 023007 023007 023007 023007 023007 023007 0230007 023007 023007 023007 023007 023007 023007 023007 0230007 0230007 0230007 0230007 0230007 0230007 0230007 0230007 0230007 0230007 0230007 0230007 0230007 0230007 02300	017622 017863 017863 017864 02106 02106 022106 022106 022107 022107 022107 023074 023074 023074 023074 023074 023074 023074 023074 023074 023074 023074 023074 023074 023074 023074 023074 02408 03408 03408	'n	148	017366				100		:
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[A] have only s < 8 significant digits, the user may compute sigrev, the corrected number of significant digits in the inverse by the formula

$$sigrev = sigdig - 8.3 + s$$

To get some idea how accurate sigrev itself was, the difference between sigrev and sigact, the actual average number of significant digits, was computed for many matrices of random numbers. When alsmo all the elements of the matrix were of the same order of magnitude, the results (for a batch of 120 matrices) were:

However, when the elements of the matrix ranged over three orders of magnitude, the results for a batch of 160 matrices were:

$$-1.77 < (sigrev - sigact) < 0.78$$

Both batches contained 100 5 by 5 matrices, the other matrices all being 10 by 10.

A DOUBLE PRECISION substitute for this subprogram is available.

I. <u>Error Indications</u> - ierror is an INTEGER variable which is used to flag errors.

ierror = -1 indicates both of the following:

- 1. Neither [A]-1 nor [X] could be computed.
- 2. [A] was either singular* or very nearly so (usually the former.

ierror = 1 indicates that no errors were detected.

Due to round-off errors, a singular matrix will rarely be flagged as singular (by ierror = -1) unless it has a row or column of zeros. However, the value of sigdig will seldom be greater than 1.5 for singular matrices.

Division by zero cannot occur in this subprogram.

No test for floating-point overflow is made in this subprogram.

^{*}When [A] is singular, a system of simultaneous equations [A][X] = [B] will have a solution if and only if [B] can be expressed as a linear combination of the columns of [A].

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Appendix C AUXILIARY LIBRARY ROUTINES

Operating descriptions and listings are presented here for LSQPØL and SID-two subroutines from the MDAC Library. These were not programmed under this contract, but are used by the FLAME code.

C. 1 LSQPØL

Least Squares Polynomial SUBROUTINE

- A. Description This subroutine subprogram computes the coefficients of the polynomial which best fits (in the least squares sense) a given function of one independent variable.
- B. Use CALL LSQPOL (x, y, w, r, n, s, m, c) where:
 - x is the REAL array, dimensioned at least n, of the values of the independent variable of the given function.
 - y is the REAL array, dimensioned at least n, of the values of the dependent variable of the given function.
 - w is the REAL array, dimensioned at least n, of the least squares weighting for each point of the given function.

 (If an unweighted least squares fit is desired, each element of this array must be a 1.)
 - r is the REAL array, dimensioned at least n, which this subprogram sets equal to the residuals at each point of the given function.
 - n is an INTEGER variable (or constant) denoting the number of points defining the given function. This number must be ≤ 50.
 - s is the REAL variable which this subprogram sets equal to the sum of the squares of the residuals.
 - m is the INTEGER variable (or constant) denoting how many coefficients the least square polynomial will have. That is, it is the degree plus one of the least squares polynomial desired. This number must be input ≤ 10.

- c is the REAL array, dimensioned at least m whose elements are set equal to the coefficients of the least squares polynomial beginning which the constant term.
- C. Support Level Supported by Programming Systems and Support Branch of Information Processing Systems.
- D. Language Used FORTRAN.
- E. Availability On FORTRAN library.
- F. Exient 40528 words.
- G Timing Not available.
- H. Restrictions None.
- I. Errol Indications The least squares problem is mathematically inconsistent aless m ≤ n, therefore, if m > n the following note is printed and execution of the program is terminated.

LSQPOL ERROR - MORE COEFFICIENTS THAN POINTS

J. Method - The method used is completely described in "Polynomial Curve Fitting With Constraint" by J. E. L. Peck in SIAM Review, volume 4, number 2, April, 1962.

This subprogram finds the ci's such that:

$$s = \sum_{k=1}^{n} w_k r_k^2$$
 is a minimum.

$$r_k = y_k - \sum_{i=1}^m c_i x_k^{i-1}$$

K. Examples - None.

C.2 SID

Single Precision Simultaneous Equation Solution and Matrix Inversion SUBROUTINE

A. Description - This subroutine subprogram will invert a REAL nonsingular square matrix, [A], and if desired, also solve linear system(s) of simultaneous equations, [A] [X] = [B], where [B] may have any number of columns. Every column of [B] must have at least one; non-zero element. The dimensions of [A] and [B] are limited only by the available core storage.

B. <u>Use</u> - CALL SID (amat, nrow, maxrow, mxcola, bmat, ncolb, mxcolb, sigdig, ierror, rtemp, itemp, scaleb) where:

1. Input

amat is a REAL two dimensional array which contains [A]. [A] will be replaced by [A]-1 during the execution of this subprogram.

nrow is an INTEGER variable or constant which denotes the number of rows in [A]. nrow may be as small as one (1) or as large as available core storage will permit.

maxrow is an INTEGER variable or constant which denotes the maximum number of rows which may be stored in the amat array. Note that a matrix may have fewer rows and/or columns than the array which contains it.

mxcola is an INTEGER variable or constant which denotes the maximum number of columns which may be stored in the amat array.

is a Real array which contains [B] if [B] is present. If [B] is present, the bmat array must have exactly maxrow rows. [B] will be replaced by [X] during the execution of this subprogram. If no [B] is present, bmat may be any variable or constant of any type.

ncolb is an INTEGER variable or constant which denotes the number of columns in [B]. If there is no [B], use ncolb = 0. ncolb may be as small as zero or as large as available core storage will permit.

mxcolb is an INTEGER variable or constant which denotes the maximum number of columns which may be stored in the bmat array. If no [B] is present, mxcolb may have any value.

2. Output

amat will contain [A]⁻¹.

bmat will contain [X] if [B] was present.

is a REAL variable which will be set equal to an estimate of the number of significant digits in the elements of the inverse. However, this estimate is based on the assumption that the elements of [A] are all accurate to eight significant digits. The Restrictions section describes a simple adjustment of sigdig which the user must make when the elements of [A] have fewer than eight significant digits.

ierror is an INTEGER variable which is used to flag errors. Its meaning is explained in Error Indications.

3. Intermediate - The following three arrays are used during calculation for the temporary storage of intermediate results. It makes no difference whether the arrays are one, two, or three dimensional. The three arrays need not all be of the same dimension.

rtemp is a REAL array which contains at least 3 nrow elements.

itemp is an INTEGER array which has at least 3. nrow elements.

scaleb is a REAL array which contains at least noolb elements.

- C. Support Level Supported by Programming Systems and Support Branch of Information Processing Systems.
- D. Language Used MSSD Standard FORTRAN.
- E. Availability On the FORTRAN library.
- F. Extent 766 words, 19 words for ALOG10, and 46 words for ALOG.
- G. Timing Not available.
- H. Restrictions This subprogram cannot find a solution to the system of simultaneous equations [A] [X] = [0].

An estimate of the average number of significant digits in the elements of the inverse is given by the argument sigdig. (See the Use section.) However, this estimate is only valid when the elements of A have eight significant digits. When the elements of

[A] have only s < 8 significant digits, the user may compute sigrev, the corrected number of significant digits in the inverse by the formula

$$sigrev = sigdig - 8.3 + s$$

To get some idea how accurate sigrev itself was, the difference between sigrev and sigact, the actual average number of significant digits, was computed for many matrices of random numbers. When alsmo all the elements of the matrix were of the same order of magnitude, the results (for a batch of 120 matrices) were:

However, when the elements of the matrix ranged over three orders of magnitude, the results for a batch of 160 matrices were:

$$-1.77 < (sigrev - sigact) < 0.78$$

Both batches contained 100 5 by 5 matrices, the other matrices all being 10 by 10.

A DOUBLE PRECISION substitute for this subprogram is available.

I. <u>Error Indications</u> - ierror is an INTEGER variable which is used to flag errors.

ierror = -1 indicates both of the following:

- 1. Neither [A]⁻¹ nor [X] could be computed.
- 2. [A] was either singular* or very nearly so (usually the former.

ierror = 1 indicates that no errors were detected.

Due to round-off errors, a singular matrix will rarely be flagged as singular (by ierror = -1) unless it has a row or column of zeros. However, the value of sigdig will seldom be greater than 1.5 for singular matrices.

Division by zero cannot occur in this subprogram.

No test for floating-point overflow is made in this subprogram.

^{*}When [A] is singular, a system of simultaneous equations [A][X] = [B] will have a solution if and only if [B] can be expressed as a linear combination of the columns of [A].

- J. Method The numerical method used is basically the Gauss-Jordan elimination with selection of maximum pivotal elements (full pivoting). A description of the Gauss-Jordan method is contained in Numerical Analysis by Kaiser S. Kunz; McGraw-Hill, 1957.
 - Special procedures are present to improve accuracy and also to minimize the frequency of overflow and underflow when the magnitudes of any of the elements in the A or B matrices are either large or small.
- K. Example The coding below will invert a matrix, solve a system of simultaneous equations, and print a warning note if the inverse has less than about one significant digit. The elements of the matrix originally in AMAT are assumed to have an average accuracy of eight significant digits.

DIMENSION AMAT (10, 10), BMAT(10, 1), RTEMP(10, 3), ITEMP(10, 3), 1 SCALEB(1)

NRQW = 6

CALL SID (AMAT, NR ϕ W, 10, 10, BMAT, 1, 1, SIGDIG, IERR ϕ R, 1 RTEMP, ITEMP, SCALEB)

IF (SIGDIG .LT. 2.) WRITE (6, 180) SIGDIG 180 FØRMAT (40H1WARNING - MATRIX INVERSE HAS ØNLY ABØUT,

1 F5. 1, 20H SIGNIFICANT DIGITS.)

```
SUBROUTINE SID (A, N, NDROW, NDCOLA, B, M, NDCOLA, SIGDIG, JERROR,
          PIVOT, INDEX, SCALEB )
C
      SID - A SINGLE PRECISION SIMULTANEOUS EQUATION SOLVER. INVERSE
                        FINDER, AND DETERMINANT SUBROUTINE
C
      DIMENSION A(NDROW, NDCOLA), B(NDROW, NDCOLB), PIVOT(N, 3),
         SCALER(M). INDEX(N.3)
      DOUBLE PRECISION DRIGP2
                    DB IGP2
      ATAG
     · / 7378697629483829.D4
C
      EPS = 1 \cdot F - 3
  712 EPS = EPS/2.
      EPSP15 = FPS + 1.5
      IF (EPSP15 .NF. 1.5) GO TO 712
      SIGMCH = ALOG10(1.522/EPS)
      RIGPW2 = DBIGP2
      PIVOT(1,1) = 0.
C
      SCALE ROWS
С
      DO 38 I=1.N
      ROWMX = 0.
       DO 28 J=1.N
       IF ((\Delta BS(\Delta(1,J))) \cdot GT \cdot ROWMX) \cdot ROWMX = \Delta BS(\Delta(1,J))
   28 CONTINUE
       IF ( ROWMX) 29, 750, 29
   29 CONTINUE
       ROWMXI = 1. / ROWMX
       100 32 J=1.N
       \Delta IJ = \Delta(I_*J)
       \Delta(1,J) = (\Delta(1,J) * ROWMXI) * BIGPW2
       IF (\Delta(I_+J)_-E0_-0_-) \Delta(I_+J) = (\Delta IJ + BIGPW2) + RDWMXI
   32 CONTINUE
       IF (M) 34, 38, 34
   34 DO 36 J=1.M
       BIJ = B(I,J)
       B(I,J) = (B(I,J) * ROWMXI) * BIGPW2
       IF (B(I,J) \cdot EO \cdot O \cdot) B(I,J) = (BIJ * BIGPW2) * ROWMXI
   36 CONTINUE
   38 PIVOT(I_{\bullet}2) = ROWMXI
       SCALE COLUMNS
С
С
       DO 10 J=1.N
       COLMX = 0.
       00 4 I=1.N
       IF (ARS(A(I,J)).GT. COLMX) COLMX = ABS(A(I,J))
     4 CONTINUE
       IF ( COLMX ) 5, 750, 5
     5 CONTINUE
       COLMXI = 1./COLMX
       DO 8 1=1.N
       AIJ = A(I,J)
       \Delta(I,J) = (\Delta(I,J) + COLMXI) + BIGPW2
       IF (A(I+J) \cdot EQ \cdot Q \cdot A(I+J) = (AIJ * BIGPW2) * COLMXI
    8 CONTINUE
   10 PIVOT(J.3) = RIGPW2 * COLMXI
```

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IF (M) 14.24.14
   14 DO 22 J=1.M
      COEMX = 0.
      DO 16 I:1,N
      IF (ABS(B(I,J)).GT. CDLMX) COLMX = ABS(B(I,J))
   16 CONTINUE
      IF (COLMX ) 17, 22, 17
   17 CONTINUE
      SCALEB(J) = COLMX / BIGPW2
      COLMXI = 1./COLMX
      DO 20 I=1 N
      BIJ = B(I,J)
      B(I,J) = (B(I,J) * COLMXI) * BIGPW2
      IF \{B(I,J) \cdot FO \cdot O \cdot\} B(I,J) = \{BIJ \neq BIGPW2\} \neq COLMXI
   20 CONTINUE
   22 CONTINUE
   24 CONTINUE
С
      INITIALIZATION
      PMONE=1.
      DO 42 J=1,N
      PIVOT(J_{\bullet}I) = 0_{\bullet}
   42 INDEX(J,3) = 0
      DO 550 I=1,N
C
Ç
      SEARCH FOR PIVOT ELEMENT
      ABPIVI=0.
   45 DO 105 J=1.N
   50 IF (INDEX(J.3)-1) 60,105,60
   60 DO 100 K=1+N
   70 IF (INDEX(K,3)-1) 80,100,80
   80 IF (ABS(A(J,K)) - ABPIVI) 100,100,85
   85 IROW=J
   90 ICULUM=K
      ABPIVI=ARS(A(J,K))
  100 CONTINUE
  105 CONTINUE
      IF (I-1) 115,120,115
  115 IF ( ABPIVI .GE. PIVMIN ) GO TO 123
  120 PIVMIN=ARPIVI
      IF (ARPIVI) 123,750,123
  123 CONTINUE
      INDEX(ICOLUM,3)=1
      PIVOTI=A(IROW.ICOLUM)
      PIVOT(I,1) = PIVOTI
C
С
      INTERCHANGE ROWS TO PUT PIVOT ELEMENT ON DIAGONAL
  130 IF (IROW-ICOLUM) 140, 260, 140
  140 PMONE =- PMONE
  150 DO 200 L=1,N
  160 SWAP=A(IROW+L)
  170 A(IROW, L) = A(ICOLUM, L)
  200 A(ICOLUM, L)=SWAP
  205 IF (M) 260, 260, 210
  210 00 250 L=1, M
  220 SWAP = B(IROW+L)
```

```
230 B(IROW_{*}L) = B(ICOLUM_{*}L)
  250 B(ICOLUM.L) = SWAP
  260 INDEX(I,1)=IROW
  270 INDEX(1,2)=ICOLUM
C
      DIVIDE PIVOT ROW BY PIVOT ELEMENT
C
      PIVINV=1.0/PIVOTI
  330 A(ICOLUM+ICOLUM) = BIGP42
  340 DO 350 L=1.N
  350 A(ICOLUM.L)= A(ICOLUM.L)*PIVINV
  355 IF (M) 380, 380, 360
  360 DO 370 L=1.M
  370 B(ICOLUM.L) = B(ICOLUM.L)*PIVINV
С
С
      REDUCE NON-PIVOT ROWS
  380 DO 550 L1=1,N
  390 IF(L1-ICOLUM) 400, 550, 400
  400 T=A(L1+ICOLUM)
      IF (T) 420,550,420
  420 A/L1.ICOLUM)=0.0
  430 DO 450 L=1.N
  450 A(11.L)=A(L1.L)-A(ICOLUM.L)+T
  455 1F(M) 550, 550, 460
  460 DO 500 L=].M
  500 B(Li*L) = B(L1*L) - B(ICOLUM*L)*T
  550 CONTINUE
С
      INTERCHANGE COLUMNS
C
С
  600 DO 710 I=1,N
  610 L=N+1-1
  620 IF (INDEX(L,1)-INDEX(L,2)) 630, 710, 630
  630 JROW=INDEX(L+1)
  640 JCOLUM=INDEX(L,2)
  650 DD 705 K=1.N
  660 SWAP=A(K,JROW)
  670 A(K, JROW) = A(K, JCOLUM)
  700 A(K+JCOLUM)=SWAP
  705 CONTINUE
  710 CONTINUE
C
      PIVOT(1,1) = PIVOT(1,1) * PMONE
С
      SIGDIG = SIGMCH - ALOGIO(RIGPW2/PIVMIN)
      IF (SIGDIG .LT. .85) SIGDIG = 0.
      UNSCALE INVERSE AND SOLUTION(S)
C
      00 720 J=1.N
      ROWMXI = PIVOT(J.2)
      00 720 I=1.N
      IF (ROWMXI .LT. 1.) GO TO 715
      A(I,J) = (A(I,J) * PIVOT(I,3)) * ROWMXI
      GO TO 720
  715 \Delta(I_{+}J) = \Delta(I_{+}J) * (PIVOT(I_{+}3) * ROWMXI)
  720 CONTINUE
     IF (M) 725, 735, 725
```

```
725 DO 730 J=1+M
ROWMXI = SCALEB(J)
DO 730 I=1+N
IF (ROWMXI *LT** 1**) GO TO 728
B(I**J) = (B(I**J) * PIVOT(I**3)) * ROWMXI
GO TO 730
728 B(I**J) = B(I**J) * (PIVOT(I**3) * ROWMXI)
730 CONTINUE
735 CONTINUE
C
IERROR = 1
RETURN
750 IERROR = -1
SIGDIG = 0*
RETURN
END
```

```
SURROUTINE LSOPUL (XI.YI.WI.RO.N.SUM.M.COEF)
     DIMENSION XI(1).YI(1).WI(1).RO(1).COFF(1)
     COMMON / ZYLSOP/
                X(50),Y(50),W(50),B(11),A(11),S(11),DSQ(11)
               ,P(50),P0(50),C(11)
     DOUBLE PRECISION X.Y.W.B.A.S.PO.P.DSO.C.G
     00\ 100\ I = 1 + N
     x(1) = x1(1)
     Y(I) = YI(I)
100 W(1) = W1(1)
     CALL ORPOLG(M-1,X,Y,W,N,B,A,S,DSQ,PO,P,G,E)
     IF(F) 1100,150,1100
150 CALL ORPOLC(M-1,C.S.A.R.G.P.PO)
     00 200 I = 1.M
200 \text{ COEF(I)} = \text{C(I)}
     SUM = 0.0
     D0 500 I = 1 + N
     K = M-1
     Z = CUEF(M)
     IF (K) 300+350+300
 300 Z = Z * XI(I) + COEF(K)
     K = K-1
     IF (K) 350,350,300
 350 \text{ RO(I)} = 7 - \text{YI(I)}
 500 \text{ SUM} = \text{SUM} + \text{RO(I)} **2
     RETURN
1100 WRITE (6,1200)
1200 FURMAT(43H1LSQPUL ERROR-MORE COEFFICIENTS THAN POINTS)
     STOP
     END
```

```
SUBRIBITINE OR POLG(N, X, Y, W, M, B, A, S, USQ, PO, P, GAMMA, ERROR)
      DOURLE PRECISION X, Y, W, B, A, PO, P , DSQ, WPP, WXPP, WYP, TEMP
      DOUBLE PRECISION WPPO.SUM.GAMMA.T
      EQUIVALENCE (TITEMPISUM)
      DIMENSION X(1),Y(1),W(1),B(1),A(1),S(1),PO(1),P (1),
                 DSQ(1)
     1
      SUM=0.
      DO 6 I=1.M
      SUM=SUM+X(I)
      MEAN=SUM/FLOAT(M)
      T=0.
      00 7 I=1.M
CD
      T=AMAX1 (T+ABS(X(I)-MEAN))
    7 \text{ TT} = DARS(X(I) - MEAN)
      IF (TT \cdot GT \cdot T) = TT
      GAMMA = T/2.
      00 8 T=1 M
      X(I) = (X(I) - MEAN)/GAMMA
  8
      ERROR=0.
      NO = M - N - 1
      IF(NO)102+103+103
      ERROR=1.
 102
      GO TO 105
 103
      B=0.
      DSO=O.
      WPP=0.
      DO 106 J=1.M
      P (J)=1.
      PO(J)=0.
      WPP=WPP+W(J)
      JF(NO)106+106+107
 107
      DSO=DSO+W(J)*Y(J)**2
      CONTINUE
      N\Delta = N+1
      DO 109 I=1.NA
      WXPP=0.
      WYP=0.
      DO 110 J=1.M
      (L)q*(L)u=qMaT
      IF(NA-I)111,111,112
      WXPP=WXPP+TEMP*X(J)*P(J)
      IF(M - I)110 + 113 + 113
 111
 113
      WYP=WYP+TEMP*Y(J)
      CONTINUE
 110
       IF(M - I)114,115,115
      S(I)=WYP/WPP
 115
      IF(NA-M)666.117.117
 114
 666
      IF(I-1)102,667,116
      DS0=DS0-S**2*WPP
 667
      GO TO 117
      DSO(I) = DSO(I-1) - S(I) **2*WPP
 116
 117
       IF(NA-1)109,109,119
      A(I)=WXPP/WPP
 119
       WPP0=WPP
       WPP=0.
       00 120 J=1.M
       TEMP = (X(J) - A(I)) *P(J) - B(I) *PO(J)
       WPP=WPP+W(J) *TEMP**2
```

```
SUBROUTINE ORPOLC (N.C.,S.A.B.G.P.PM)
     DOUBLE PRECISION C.S.A.B.G.P.PM.T1.T2
     DIMENSION C(1),S(1),A(1),B(1),P(1),PM(1)
     M = N + 1
     00 300 L=1.M
     C(L)=0.
     PM(L)=0.
300 P(L)=0.
     P=1.
     C = S
     IF (N) 301,303,301
 301 00 302 L=1.N
     T2=0.
     J=L+1
     DO 302 M=1.J
     T1 = (T2-A(L) *P(M)-B(L) *PM(M))/G
     T2 = P(M)
     PM(M) = P(M)
     P(M)=T1
302 C(M)=C(M)+T1 #S(L+1)
 303 RETURN
     END
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13 ABSTRACT							
As part of the Optical Signatures Program, McDonnell Douglas Astronautics							
Company-West has developed the initial working model of a code to describe							
the gross features of rocket-plume radiation for altitudes above 75 n mi.							
The main effort is the construction of a scheme for integration of an							
arbitrary function through an arbitrary axisymmetric rocket plume, with any							
specified look angle, plume direction, and vehicle velocity direction.							
Radiances are presented as integrated values in a specified spectral band.							
The equations used and a printout of the code and of a sample application are							
included.							

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